

# Tools and Monte-Carlos

for the

**New Physics** 

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- What's a tool and what it takes to make one
- New Physics vs the Standard Model Physics
- Modular Structures and interfaces





# **OUTLINE**

A. From the parton to the detector not necessarily BSM

B. From the Lagrangian to the MEG codes talk, interface

**NLO** issues

not necessarily BSM

Tools for Dark Matter
Definitely BSM

#### **OUTLINE A and B**

- What's a tool and what it takes to make one: Structure of an event
- Components of a MC EG (Monte Carlo Event Generator)
- Integration and MC techniques (probably skip)
- PS: Parton Shower in a MC
- Matrix Element vs PS
- ME generation and ME generators
- Modular structure of codes, Les Houches Accords
- Tools for the New Physics

#### Further Reading and from where I borrowed

- Frank Krauss Bonn Lectures, 2006

  http://projects.hepforge.org/sherpa/dokuwiki/publications/presentations/index
- Fabio MALTONI HEPTOOLS School, Torino, 2008 http://personalpages.to.infn.it/ maina/scuola08/Maltoni\_Torino08.pdf
- Steve Mrenna CTEQSS, CTEQ05
- Peter Richardson CTEQ06 School, IPPP Durham, 2006
- Mike Seymour CERN Training Lectures 2003 http://seymour.home.cern.ch/seymour/slides/CERNlecture1.ppt
- Torbjrn Sjostrand, 2006 European School of HEP, Aronsborg YETI06, IPPP Durham, see Pythia website http://www.thep.lu.se/ torbjorn
- Brian Webber 1st MCnet School, IPPP Durham 2007
- Les Houches Guidebook Les Houches Guidebook to Monte Carlo Generators for Hadron Collider Physics, hep-ph/0403045
- R.K. Ellis, W.J. Stirling and B.R. Webber QCD and Collider Physics

## Further Reading and where you can find info on (and download) the codes

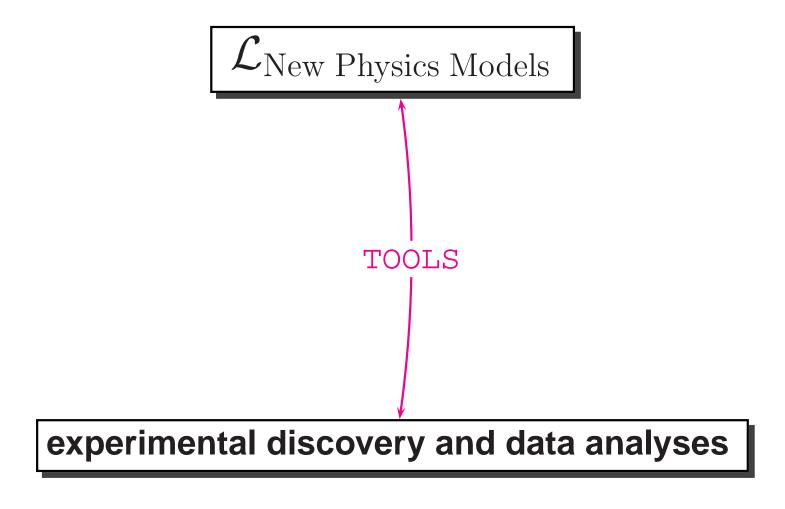
- Les Houches Guidebook Les Houches Guidebook to Monte Carlo Generators for Hadron Collider Physics, hep-ph/0403045
- BSM Tools Repository: http://www.ippp.dur.ac.uk/montecarlo/BSM/.
- SUSY Bestiary of Public Codes: Allanach, arXiv:0805.2088 [hep-ph].
- SUSY Tools for Dark Matter and at the Colliders. Fawzi Boudjema, Joakim Edsjo, Paolo Gondolo, arXiv:1003.4748
- My webpage
- Les Houches May 30-Jun 17, 2010 (not just a tools Workshop....)
- MC4BSM, Tools for SUSY and the New Physics Series (Google them)

#### **Nobel Dreams**

Great Idea: A New Physics Model

**FINAL AIM** 

Nobel Prize if LHC validates!



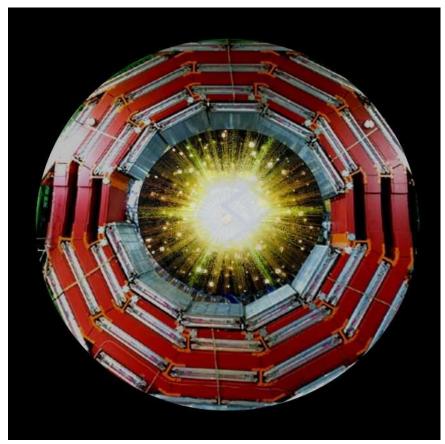
# For a long time, Tools and MC



Before LHC started

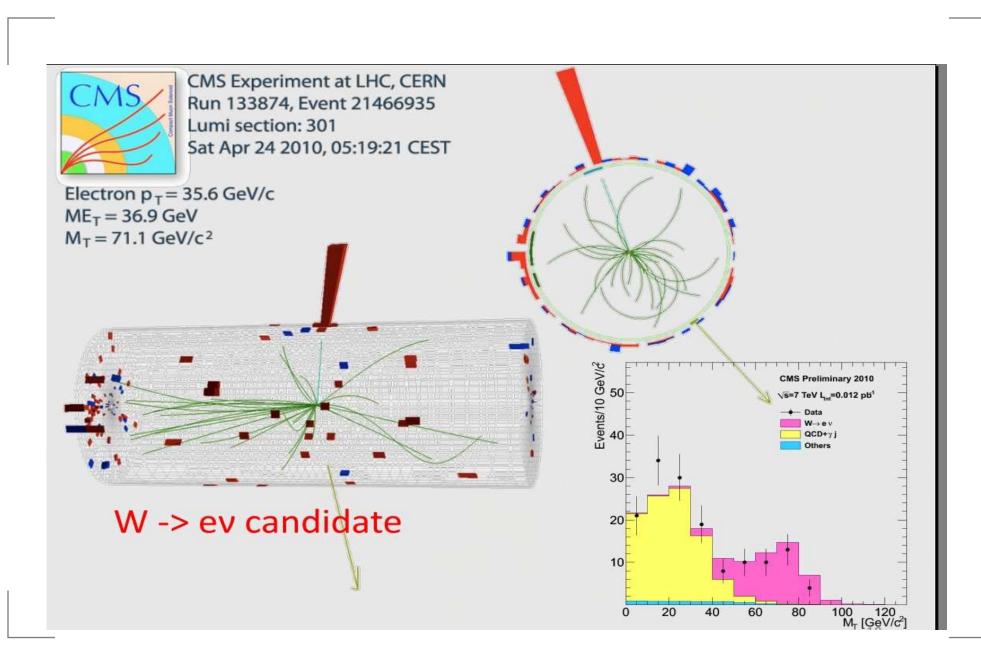
# For a long time, Tools and MC



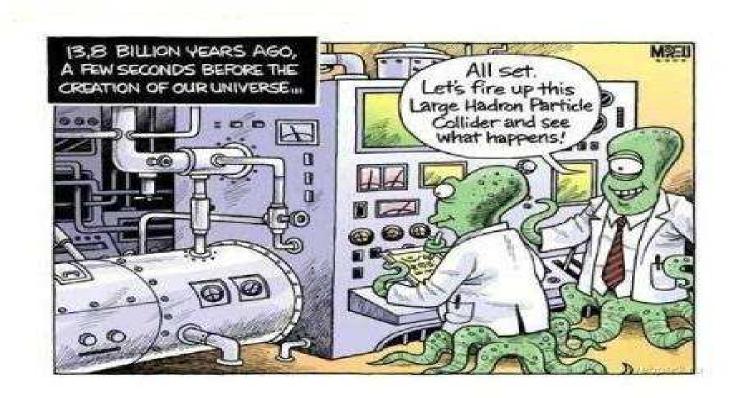


is the Higgs in there? or any other New Physics

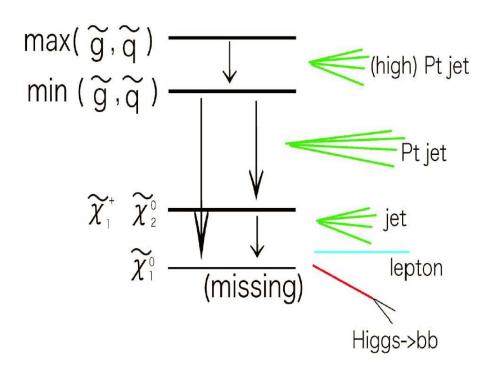
## Data from LHC, after some cleaning



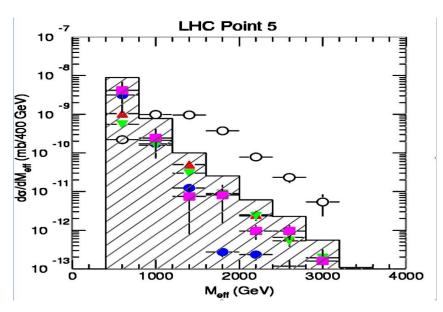
## Turn on the machine!



## in 1998 we were told to expect an early SUSY discovery

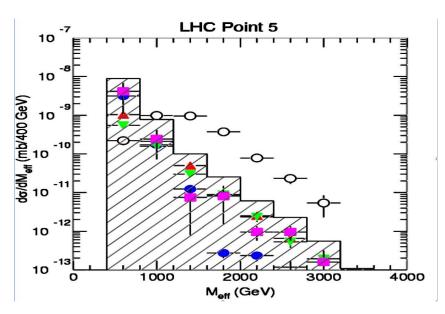


# ATLAS TDR (same with CMS)

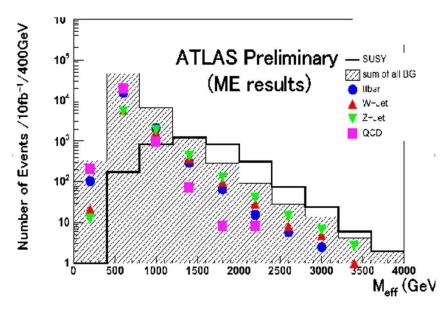


ATLAS TDR 98 (mSUGRA point, PreWMAP)

## ATLAS TDR (same with CMS)

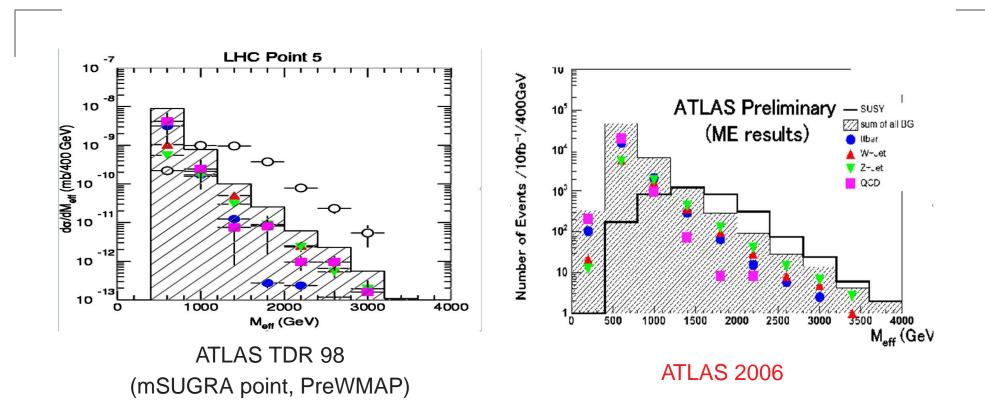


ATLAS TDR 98 (mSUGRA point, PreWMAP)



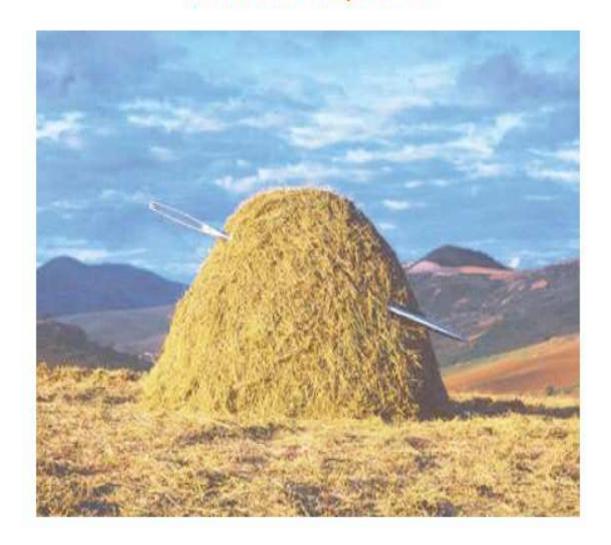
**ATLAS 2006** 

## ATLAS TDR (same with CMS)

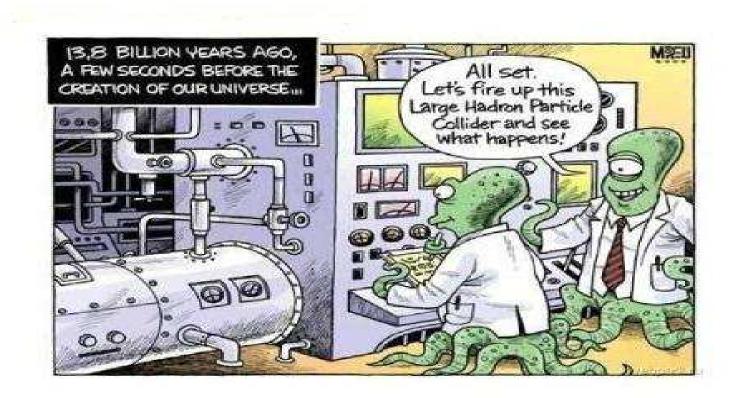


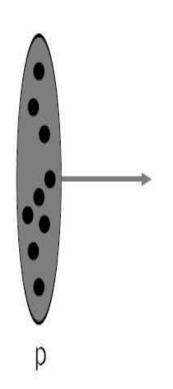
# What happened?

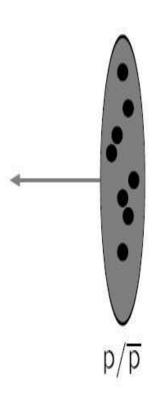
# What we hope for!



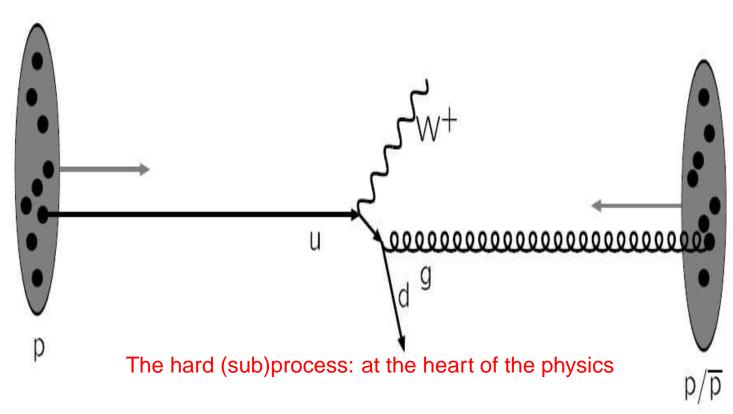
## Let's turn the machine again, slow motion!



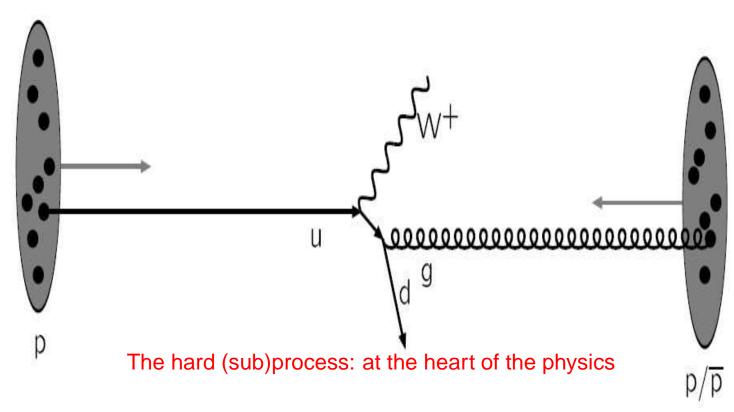




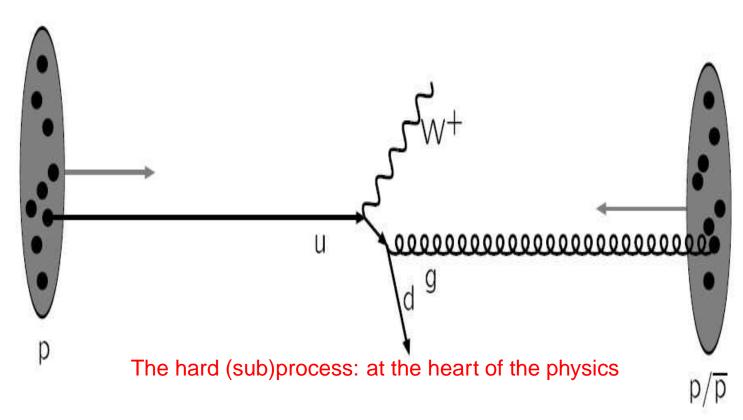
Incoming beams: partons densities



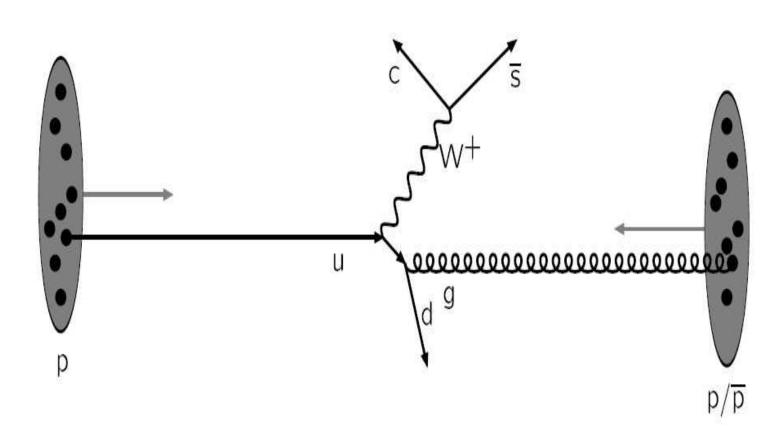
Hard process is well understood and well described: relies on a firm perturbative framework.



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- described by Matrix Elements (ME)
  This does not mean that it is very well calculated

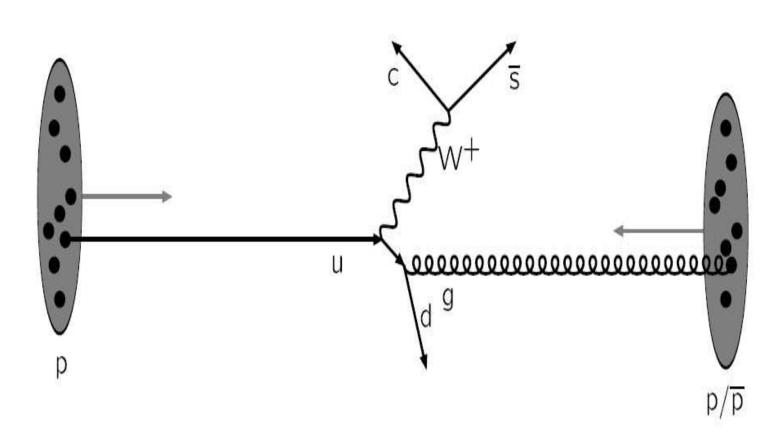


- Hard process is well understood and well described: relies on a firm perturbative framework.
- described by Matrix Elements (ME)
  This does not mean that it is very well calculated
- issue of higher order (NLO), most calculations only LO say



Decays of resonances: correlated with hard process

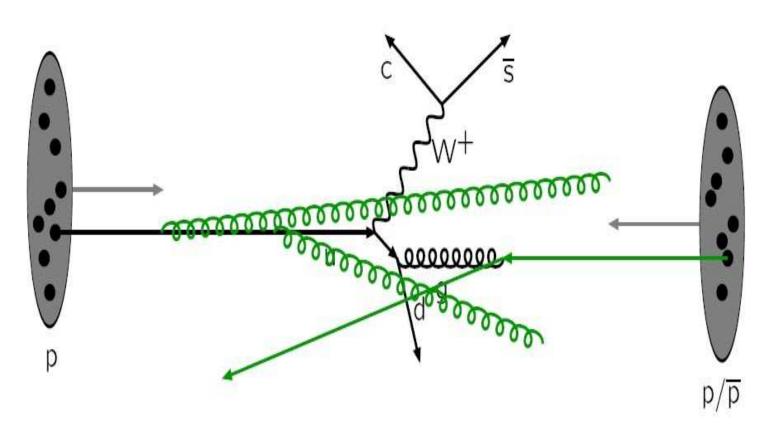
ullet Approximation: W on-shell



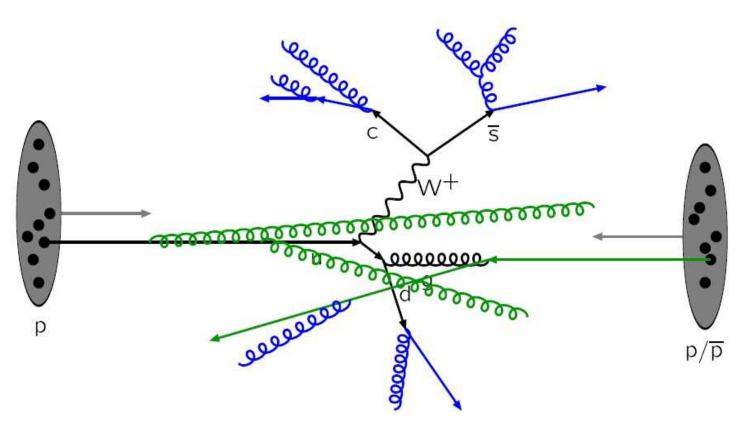
Decays of resonances: correlated with hard process

• Approximation: W on-shell

Spin effect in decays?

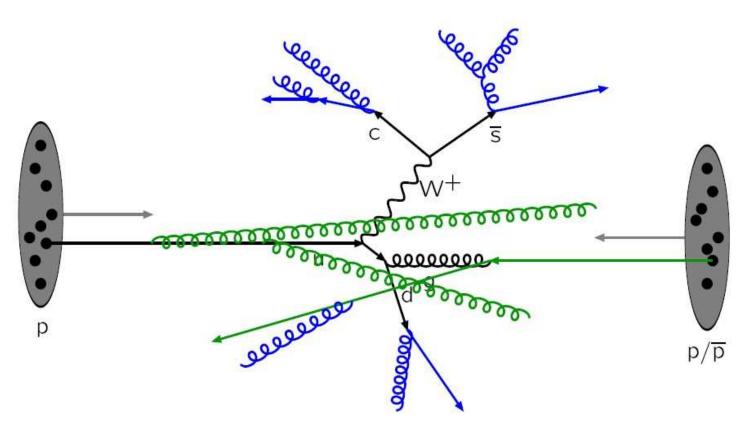


ISR: Initial State Radiation
Space-like parton showers (PS)



FSR: Final State Radiation

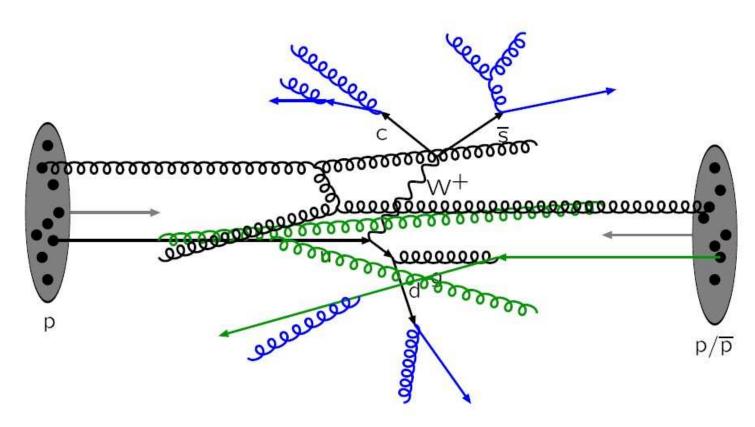
time-like parton showers (PS)



FSR: Final State Radiation

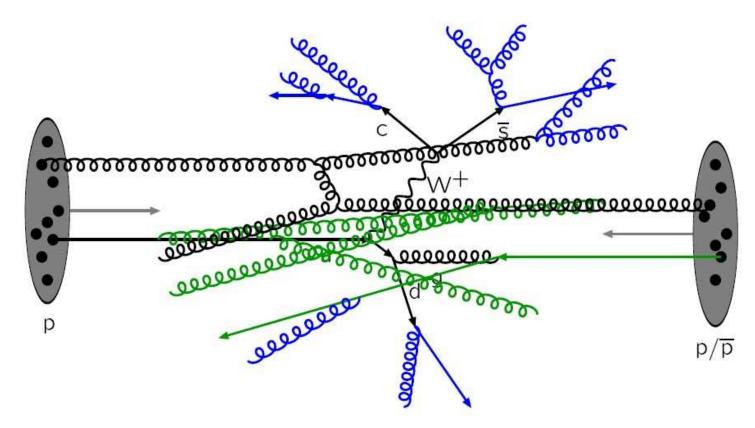
## time-like parton showers (PS)





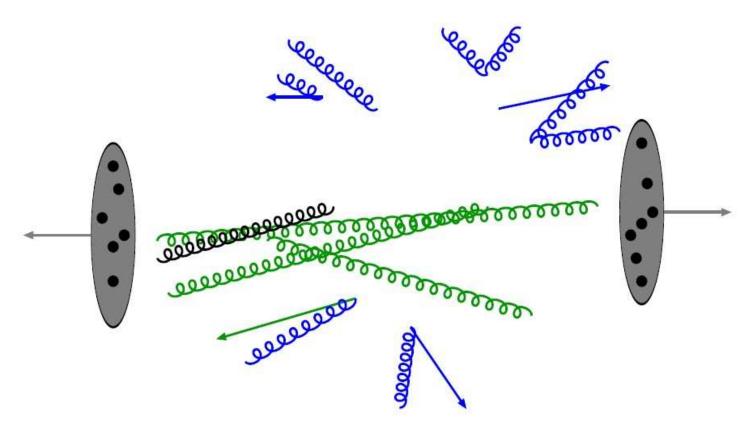
Multiple parton-parton interactions (MPI)

The muck



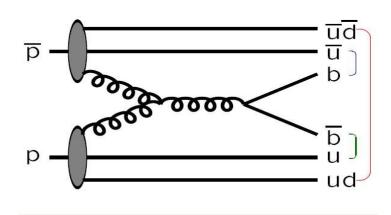
MPI with ISR and FSR!

The muck

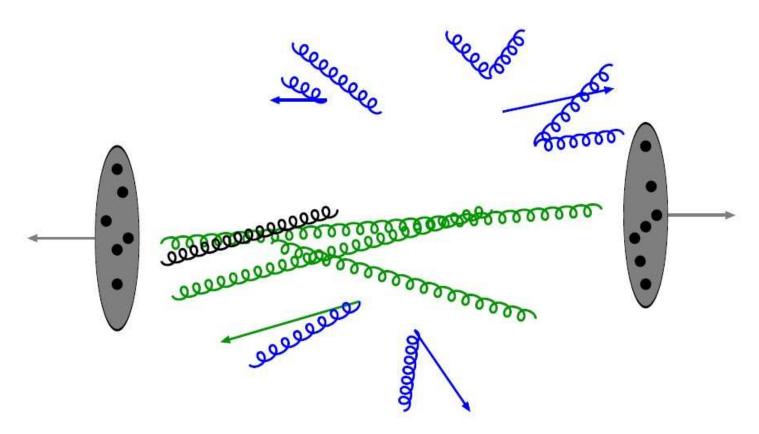


Beam remnants and other outgoing partons!

The muck

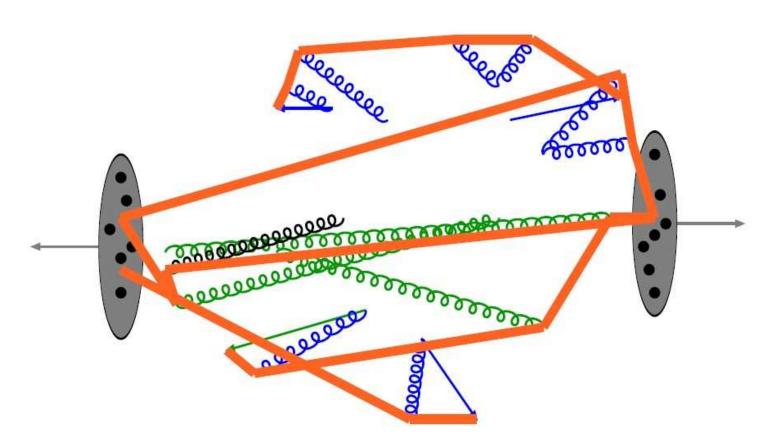


Beam remnants: coloured remains of the proton not taking part in the hard process, but they are colour connected to the hard process.

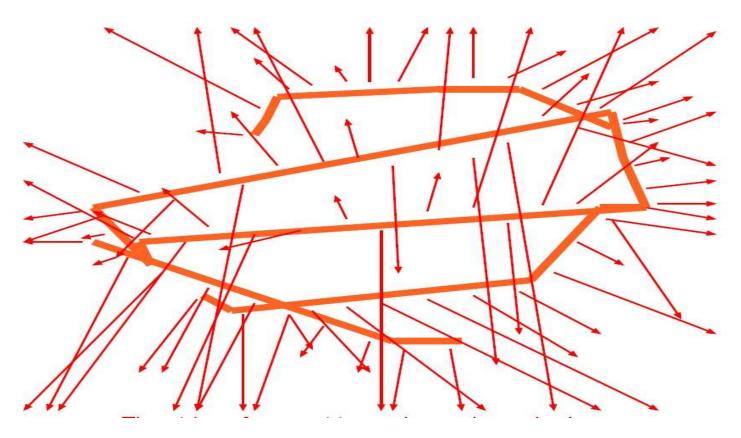


Beam remnants and other outgoing partons!

The muck: UE is not understood. It is modelled through some data that do not give strong constraints and hence extrapolation can be doubted.

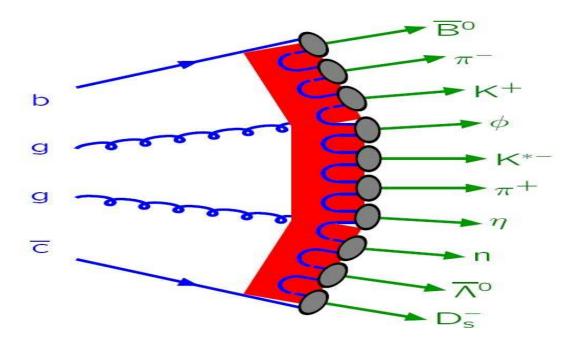


Everything is connected by colour confinement (here strings)



The strings fragment to produce hadrons

Movie: The structure of an event, hadronisation

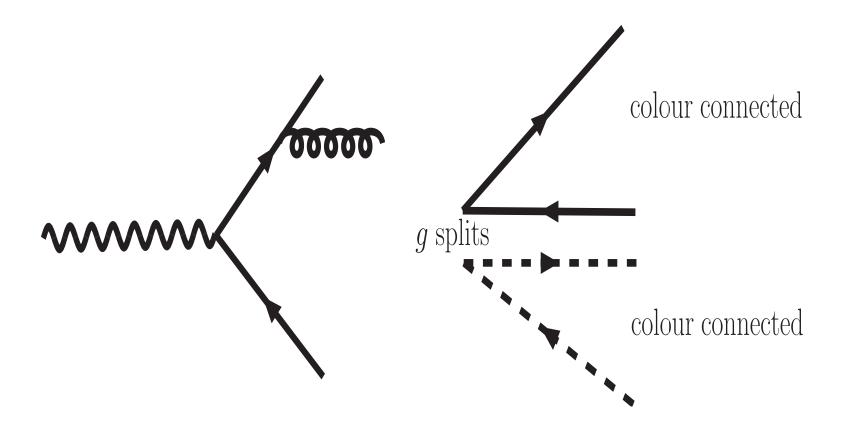


Hadronisation: Clusters to produce hadrons (Cluster Model, HERWIG)

Hadronisation and colour connection assumes standard QCD. In some exotic models (Rp violation) nor so trivial to perform this step.

## hadronisation, HERWIG

## after angular ordering cluster as

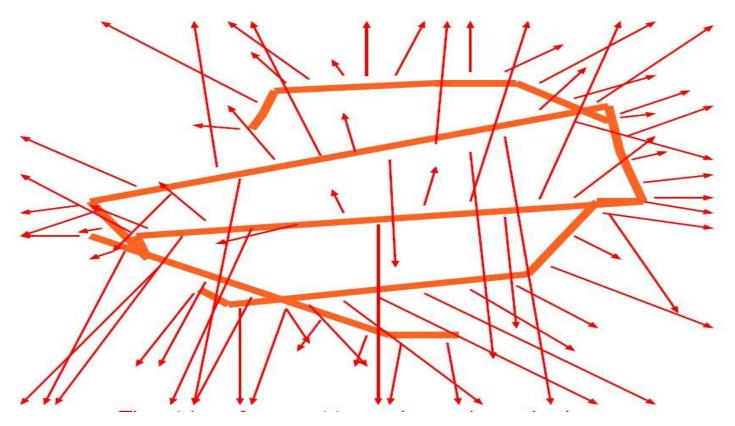


## New colour structures $R_p$ violation

$$\mathbf{W}_{R_p \text{viol}} = \dots + \frac{1}{2} \lambda_{IJK}^{"} \boldsymbol{\epsilon^{c_1 c_2 c_3}} \bar{U}_{c_1}^I \bar{D}_{c_2}^J \bar{D}_{c_3}^K \dots$$

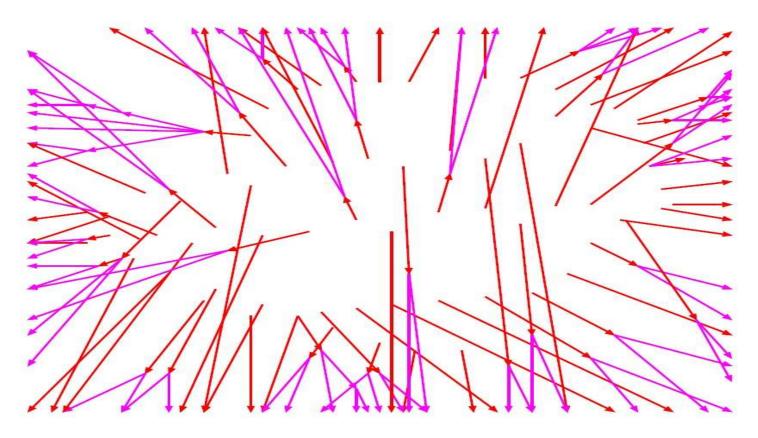
Leads to  $\tilde{q} \rightarrow qq$  baryon number violation with very exotic colour flow.

You will need a PS/EVG expert!

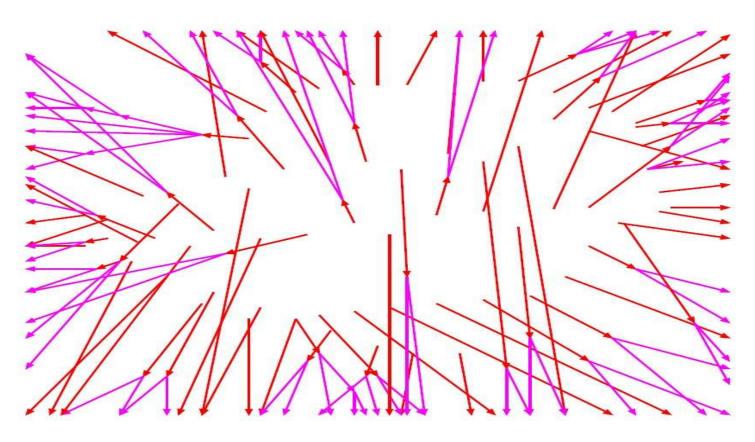


The strings fragments to produce hadrons (strings model)

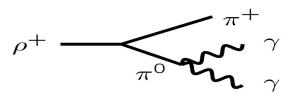
Hadronisation is not really calculated from first principles, however it is modelled through various data and hence it is considered reliable

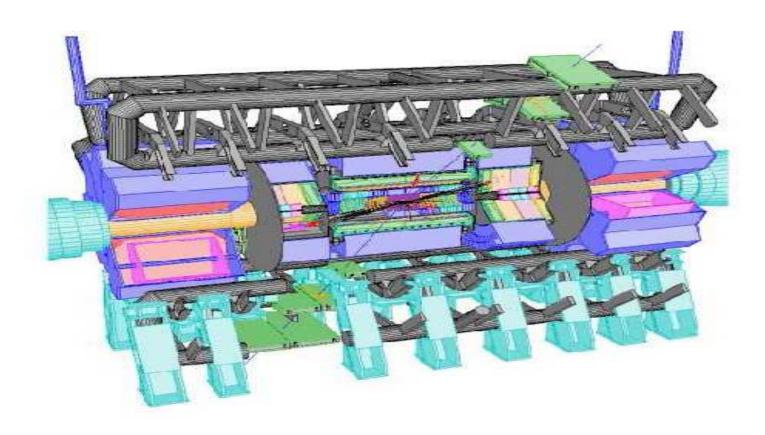


Hadrons decay



# Hadrons decay





These are the particles that hit the detector

#### Parts of a MC EG

- Parton Shower is well understood, perturbation theory with a few approximations
- Hadronisation is not really calculated from first principles, however it is modelled through various data and hence it is considered reliable
- UE is not understood. It is modelled through some data that do not give strong constraints and hence extrapolation can be doubted.(minimum bias??)
- Important to have a "clear" picture of the physical situation

### MC is probabilistic, divide and conquer

generate events with as much details as possible:

```
W will decay. To 	au?, 	au will decay, there is no quark, only hadrons,... production comes with non negligible radiation
```

- $m{ ilde{ heta}}_{ ext{final state}} = \sigma_{ ext{hard process}} \; \mathcal{P}_{ ext{tot}}$
- lacksquare Divide and Conquer : each  $\mathcal{P}_i$  handled in turn
- an event with n particles involve about 10n random choices (flavour, mass, momentum,spin,...). At the LHC expect about 100 charged and 200 neutral particles, thus totalling a few thousand choices

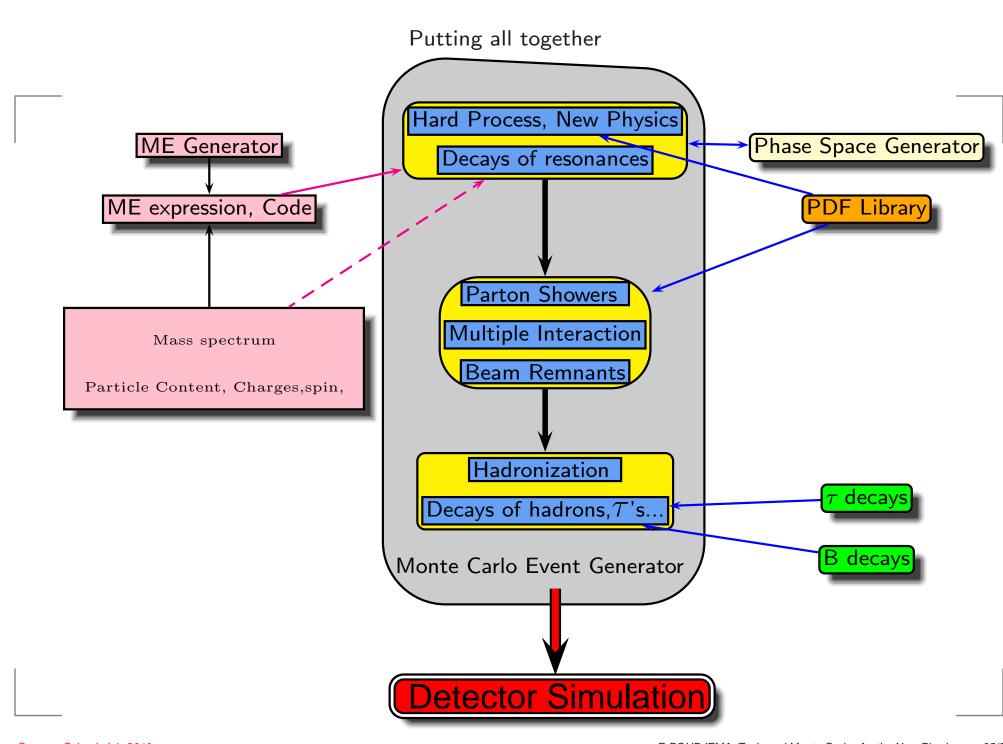
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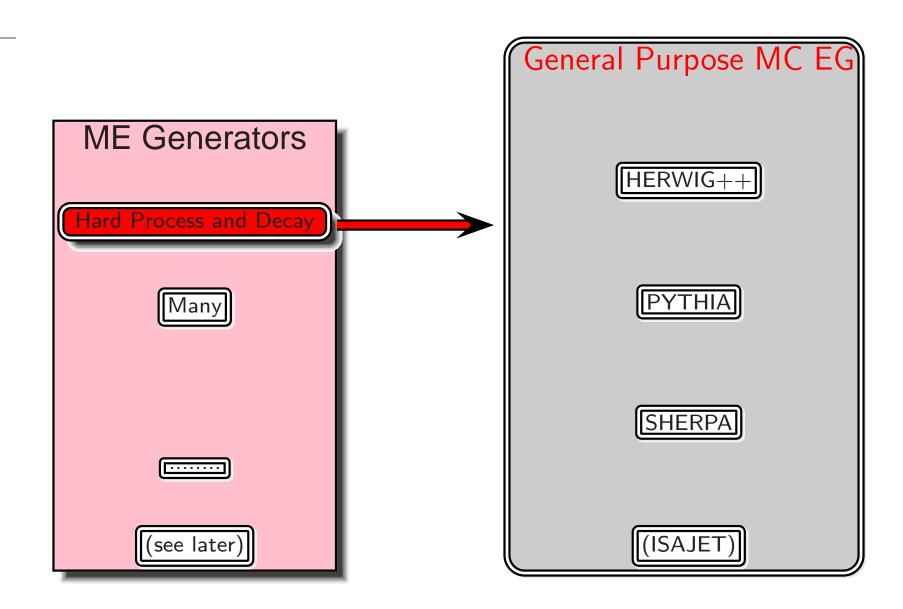
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Divide and Conquer : each  $\mathcal{P}_i$  handled in turn  $\to$  Modular Structure



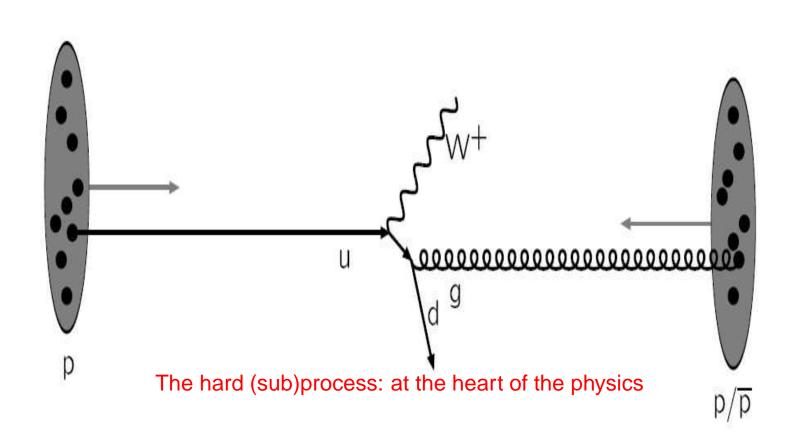


### The Hard Process

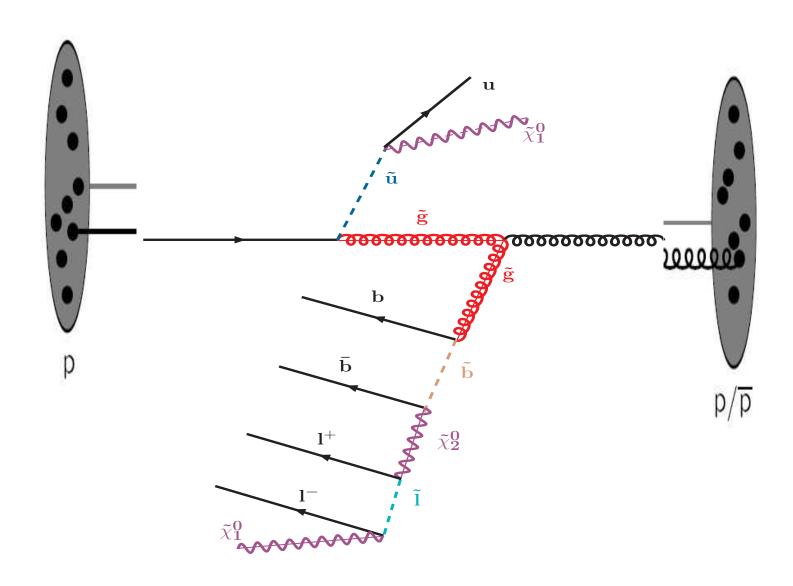
# The Hard Process and MEG

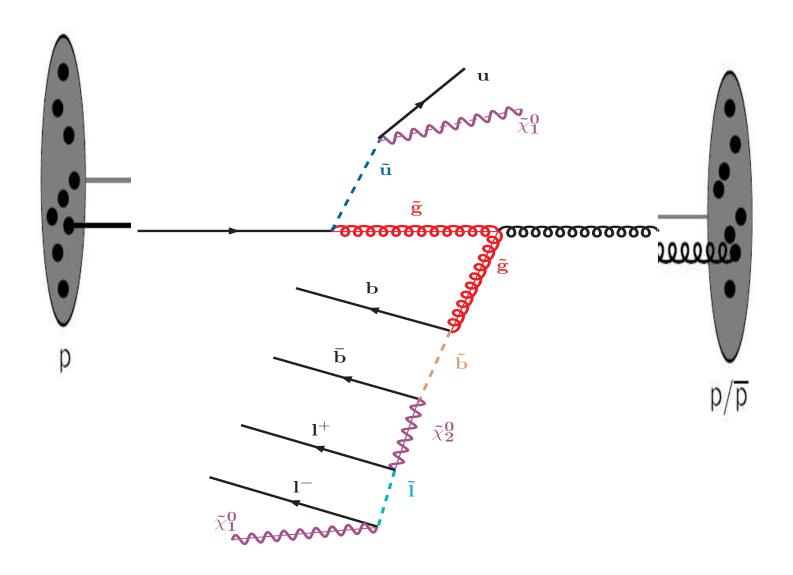
- 1. Get the partonic matrix element or matrix element squared
- 2. convolute over the PDF
- 3. integrate over phase space
- 4. 2 and 3 means integration which requires numerical evaluation  $\rightsquigarrow$  MC techniques

## The Hard Process

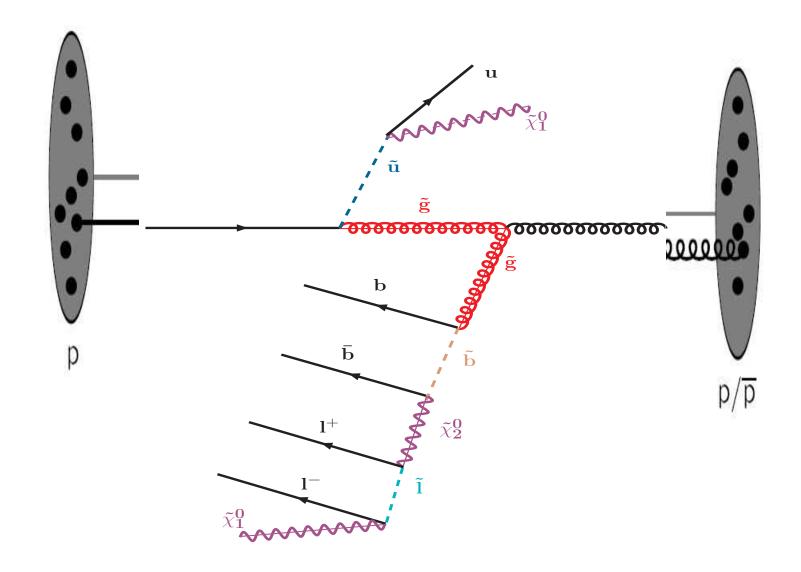


# The Hard (sub)Process at the heart of the physics

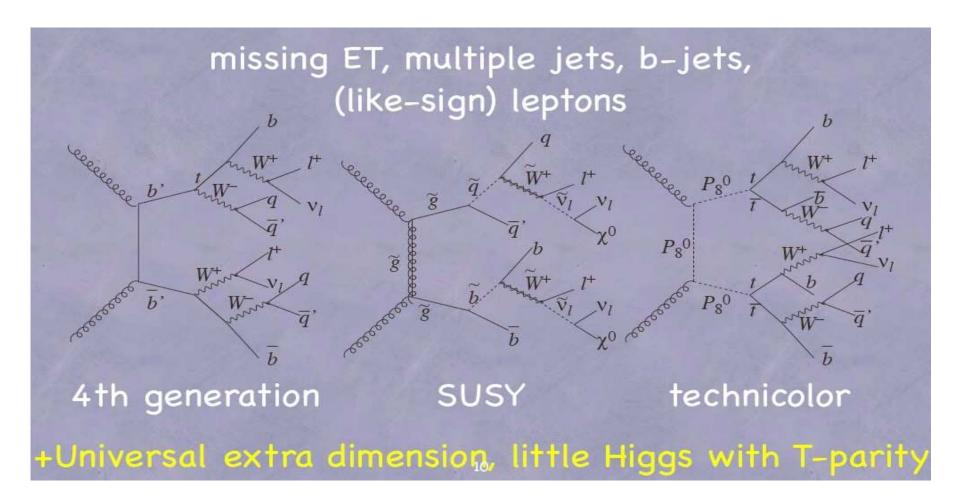




$$ug 
ightarrow ilde{u} ilde{g}$$
 or?



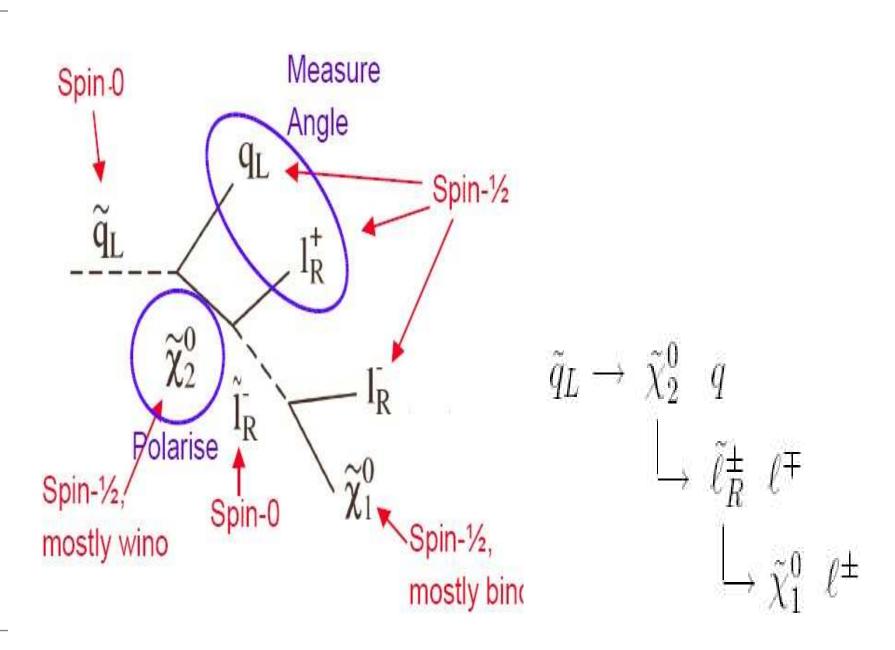
 $2 \to 7$  process, successive decays, huge number of diagrams,...



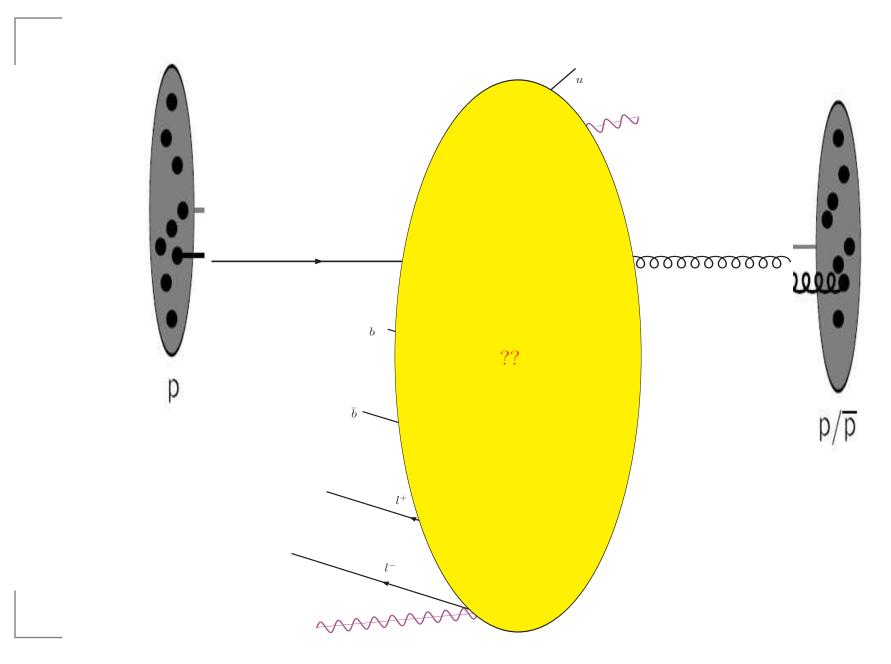
Is it necessarily DM candidate?

stable at the scale of LHC detectors, 1ms, not age of the Universe....

Next step: Properties of DM as an example, details of couplings and spin assignments.



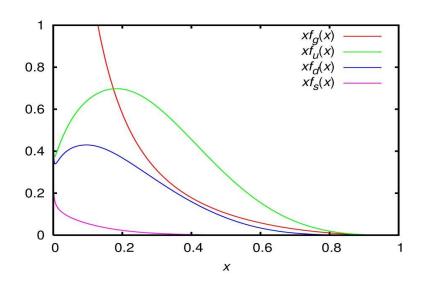
# Integration: PDF and Cross sections



### Factorisation and Parton Distribution Functions

$$\sigma_{pp\to X} = \sum_{a,b=q,\bar{q},g} \int dx_1 dx_2 f_a(x_1,\mu^2) f(x_2,\mu^2) \hat{\sigma}_{ab\to X}(\hat{s},\mu^2)$$

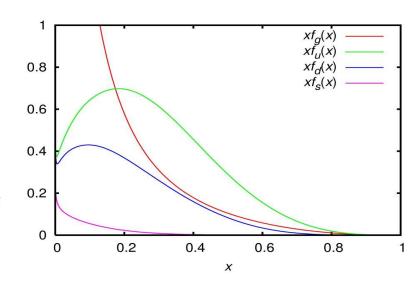
 $f_i(x,\mu^2)$  is the Parton Distributions Function  $\mu^2$  is the factorisation scale ! Many libraries exist (CTEQ, MRSx) reliable in the range  $10^{-3} < x < 0.8 \ (2 {\rm GeV})^2 < \mu^2 < (1 TeV)^2$ 



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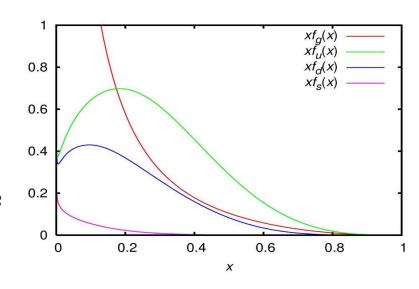
Phase Space

$$\hat{\sigma}_{ab\to X} = \frac{1}{2\hat{s}} \sum_{spin...} \int_{\Phi_N} |\mathcal{M}|^2 d\Phi_N$$

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Phase Space

$$\hat{\sigma}_{ab\to X} = \frac{1}{2\hat{s}} \sum_{spin,...} \int_{\Phi_N} |\mathcal{M}|^2 d\Phi_N$$

## MEG, Monte-Carlo and Integration

# Matrix Elements Generators

Monte Carlo as Integrator

### Monte-Carlo and Integration

At the heart of the ME is the hard process, that is where the physics lies and that is what gives the probability of a particular event For the hard process

- **amplitude**  $\mathcal{M} \longrightarrow |\mathcal{M}|^2$  first part of what an MEG should get (the dynamics)
- $N_{\rm evt,cuts} \propto \int d\sigma = \int |\mathcal{M}|^2 d\Phi(n)$
- Integration over a phase space with of large number n of dimensions, each particle  $\rightarrow 3$  variables (momenta)
- $ightharpoonup Dim[d\Phi(n)] \sim 3n$

$$d\Phi(n) = \left(\prod_{i} n \frac{d^2 p_i}{(2\pi)^3 (2E_i)}\right) (2\pi)^4 \delta\left(P_{in} - \sum_{i}^{n} p_i\right)$$

different MEG implement different techniques for the integration.

#### Monte-Carlo Definition

- MC is a numerical method for calculating/estimating an integral based on a random evaluation of the integrand
- Particularly useful because one deals with a large number of (integration) variables (momenta of particles)
- Limits of integration (cuts) are often complicated
- Integrand is a convolution of different functions

### Summary of what you should have seen: Peter Skands and Tim Stelzer

- MC integration through random number generation
- ullet MC converges as 1/sqrtN for any d-dim integral
- Trapezium, Simpson better but only for d-dim small, not the case of HEP
- Importance sampling (change of variables)
- Importance + Stratified sampling (VEGAS/BASES)
- Multi-channel
- event weight, event generation

No new NEW PHYSICS in here

JUMP ▶

### One dimension, example

$$I = \int_{x_1}^{x_2} f(x)dx = (x_2 - x_1) < f(x) >$$
 (usually  $x_1 = 0, x_2 = 1$ )

The average can be calculated by selecting N values  $randomly x_i, i = 1, \cdots N$  from uniform distribution, calculate  $f(x_i)$ 

$$I = I_N = \frac{1}{N}(x_2 - x_1) \sum_{i=1}^{i=N} f(x_i) = \frac{1}{N} \sum_{i=1}^{i=N} W(x_i)$$
  $W(x_i) =$ weight

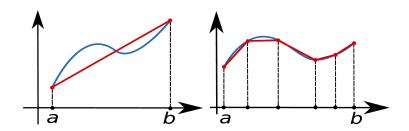
- Sum is invariant under reordering ( randomize)
- lacksquare Obviously approximation better if number of points N is larger
- Error given by the Central Limit Theorem

$$I \simeq I_N \pm \sqrt{\frac{V_N}{N}} = I_N \pm \sqrt{\frac{\langle f^2 \rangle - \langle f \rangle^2}{N}}$$

**●** MC converges as  $1/\sqrt{N}$ 

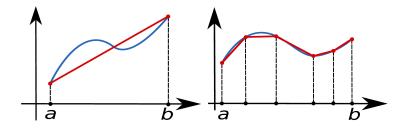
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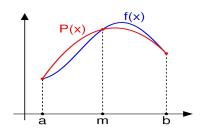


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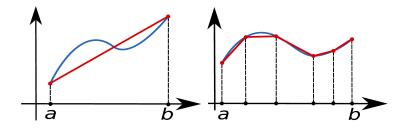


Simpson (quadratic interpolation)  $\propto 1/N^4$  (if derivative exists)

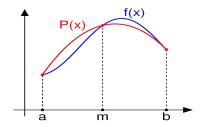


$$I \simeq I_N \pm \sqrt{\frac{V_N}{N}} = I_N \pm \sqrt{\frac{\langle f^2 \rangle - \langle f \rangle^2}{N}}$$

- **●** MC converges as  $1/\sqrt{N}$
- ightharpoonup compare to trapezium rule convergence  $\propto 1/N^2$  (if derivative exists)



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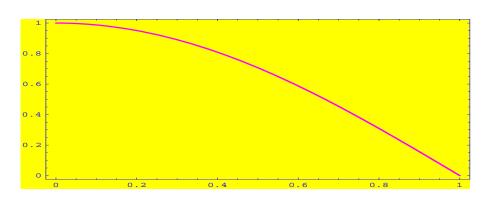
but this is only in one dimension!

- Convergence may seem slow  $\sqrt{1/N}$ , but it can be estimated easily
- MC error does not depend on # of dimensions, d,  $\propto 1/\sqrt{N}$  Trapeze  $\propto \to 1/N^{2/d}$  Simpson  $\propto \to 1/N^{4/d}$
- ${\color{red} \bullet}$  in MC one can improve convergence by minimising  $V_N$  while keeping the same number of points N
- Importance Sampling: non uniform sampling more efficient
- Convergence improved by putting more samples in regions where function is largest (where variance is largest)
- ullet Hint: observe that if f(x)=cste then  $V_N=0$   $\to$  make f as a close to a constant as possible!

$$I \simeq I_N \pm \sqrt{\frac{V_N}{N}} = I_N \pm \sqrt{\frac{\langle f^2 \rangle - \langle f \rangle^2}{N}}$$

## **Example: Importance Sampling**

Take 
$$f(x)=cos\pi x/2$$
 then  $I=2/\pi=0.637$  MC,  $I_N=0.637\pm \frac{0.308}{\sqrt{N}}$  (0.308 =  $\sqrt{V_N}=\sqrt{1/2-(2/\pi)^2}$ )

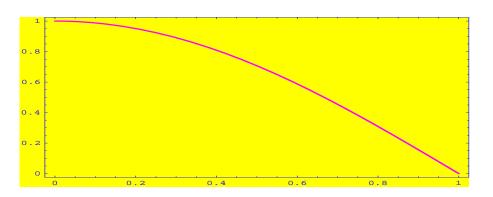


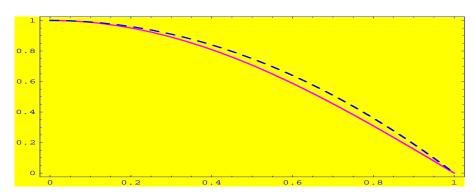
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$$= \int_{y_1}^{y_2} dy \frac{\cos \pi x[y]/2}{1 - x[y]^2}$$

MC, 
$$I_N = 0.637 \pm 0.031/\sqrt{N}$$

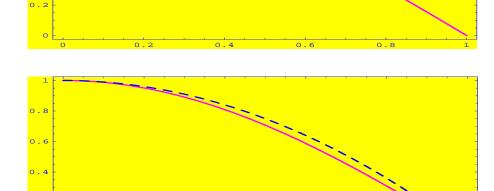




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MC, 
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- **●** For the same accuracy  $N \rightarrow N/100$  events
- We have in fact made a change of variables
- Note however that change of variables may be not so trivial and requires that one knows the function, here is relatively ok

$$y = x - x^3/3!$$

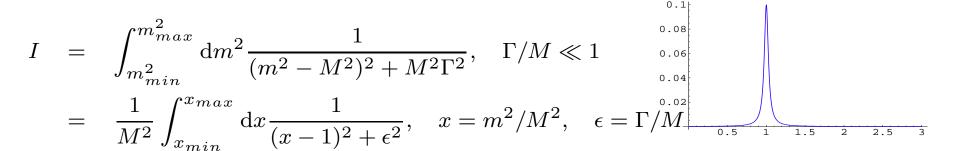
### Over a Breit-Wigner distribution

in HEP many sharp peaks from resonances, apart from peaks due to forward scattering,...

$$I = \int_{m_{min}^2}^{m_{max}^2} dm^2 \frac{1}{(m^2 - M^2)^2 + M^2 \Gamma^2}, \quad \Gamma/M \ll 1$$
$$= \frac{1}{M^2} \int_{x_{min}}^{x_{max}} dx \frac{1}{(x - 1)^2 + \epsilon^2}, \quad x = m^2/M^2, \quad \epsilon = \Gamma/M$$

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change of variable  $x = \varepsilon \tan \theta + 1, dx = \varepsilon (1 + \tan^2 \theta) dt$ 

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The peak has been smoothed out completely.

Flat distribution, the error has been reduced to 0

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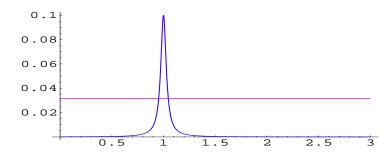
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### Non-uniform, importance sampling

Unfortunately we can not always do the Jacobian trick efficiently, we do not always know  $f(\boldsymbol{x})$ 

However, as we have seen, finding a simple function, p(x), that approximate f(x) reduces the error drastically (up to normalisation) take

$$p(x), \int_{x_1}^{x_2} p(x) = 1, \qquad \to I \qquad = \qquad \int_{x_1}^{x_2} dx p(x) \frac{f(x)}{p(x)}$$

$$I \qquad = \qquad \left\langle \frac{f}{p} \right\rangle \pm \frac{1}{\sqrt{N}} \sqrt{\left\langle \frac{f^2}{p^2} \right\rangle - \left\langle \frac{f}{p} \right\rangle^2}$$

Sample according to p(x) and make f/p as small as possible.

## VEGAS (BASES) Importance+Stratified Sampling

Unfortunately we usually do not know much about f(x)

But as we sample we can know more, reconstruct p(x) piecemeal, with step function

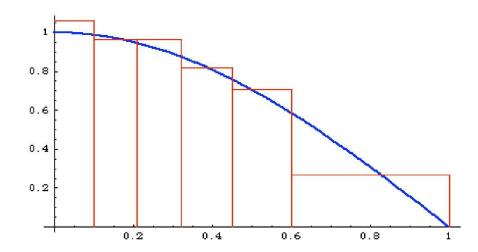
$$p(x) = \frac{1}{N_b} \Delta x_i \text{ for } x_i - \Delta x_i \le x \le x_i$$

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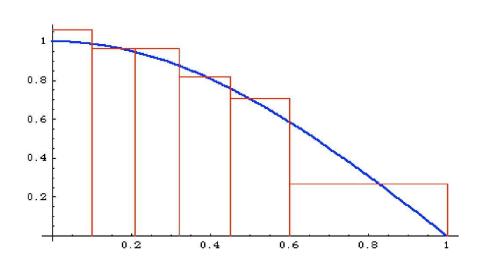


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- Improve the fit by generating more points where f(x) is large, *i.e* where the variance is large
- Adjust the bin size so that each bin has the same area

Iterative algorithm: VEGAS

The approach can be directly generalised to d dimensions if one can write the factorised from  $p(\vec{x}) = p(x) \times p(y) \times \cdots$ 

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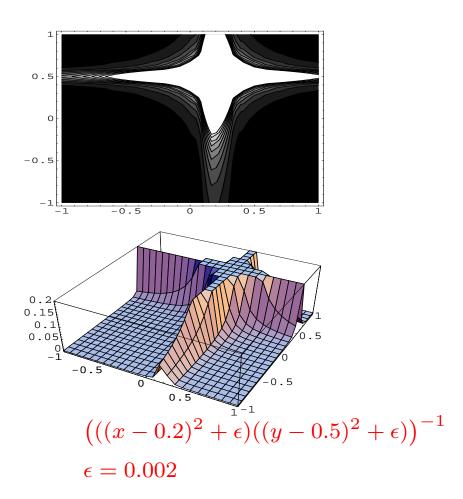
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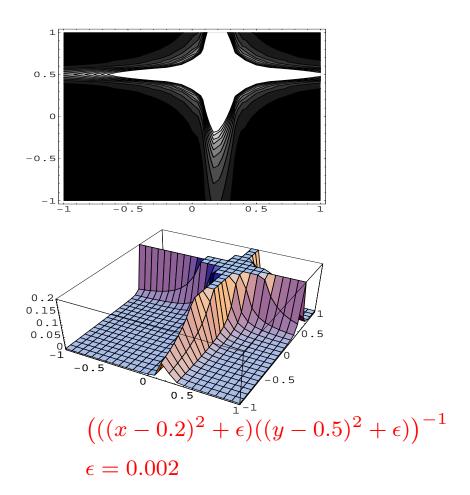
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- A scattering amplitude may have many peaks each aligned on a different invariant

# VEGAS and alignment

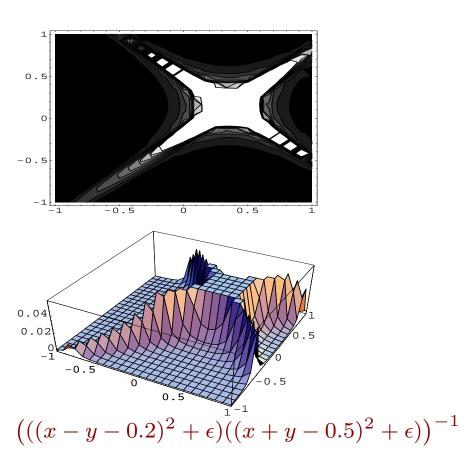


Ok for VEGAS, peaks aligned along the axes

# VEGAS and alignment

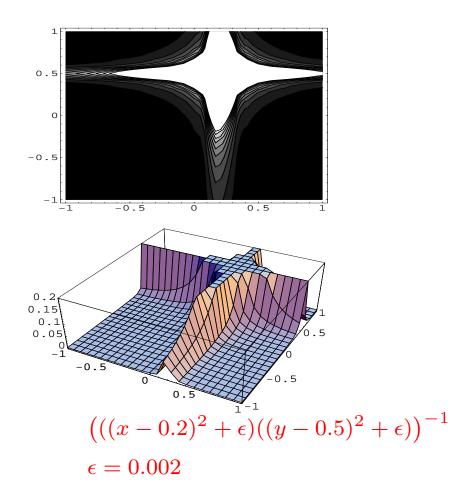


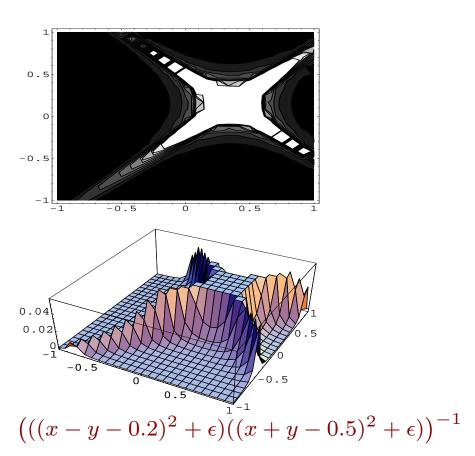
Ok for VEGAS, peaks aligned along the axes



NOT Ok for VEGAS, here rotate the axes

# VEGAS and alignment





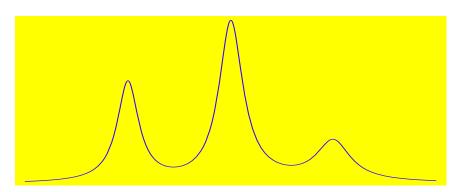
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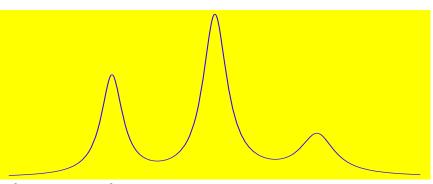
For physical processes we usually know where the peaks are

Suppose instead of having one peak we have an integral with lots of peaks, say from the inclusion of excited  $\rho$  resonances in some process.

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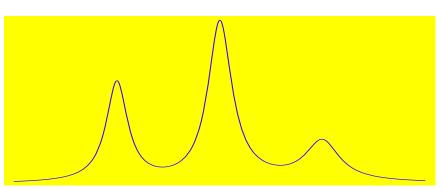


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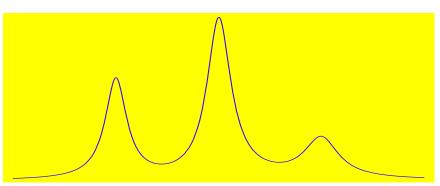
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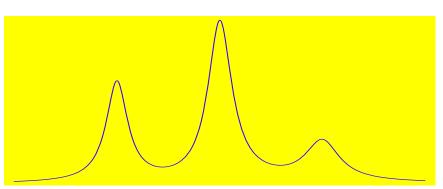
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Pick one of the integrals (channels) with prob  $\alpha_i$  then calc. weight  $\alpha_i$  can be automatised.  $\int$  does not depend on  $\alpha_i$  but  $V_N$  does

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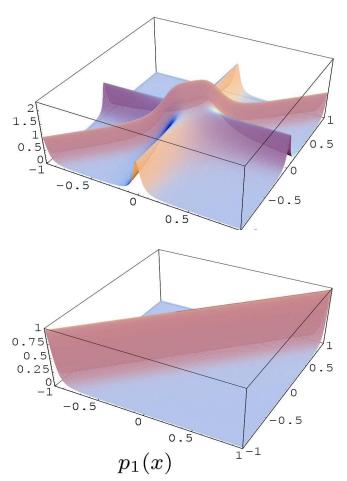
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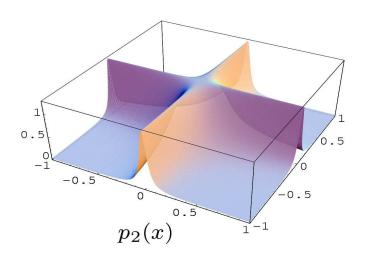
In general not each channel is invertible  $\rightarrow g_i$  (peaking may be more complicated) N coupled equations for  $\alpha_i$ , so best when number of channels small.

This is the method (multi-channel) used in the most sophisticated codes.

# Multichannel, many dimensions



- what to do here?
- decompose into different channels



For physical processes we usually know where the peaks are

#### cross section integrator vs event generator

$$d\sigma(u\bar{u} \to Z^0 \to d\bar{d}) = \frac{1}{\hat{s}} |\mathcal{M}|^2 \frac{d\cos\theta d\phi}{8(2\pi)^2}$$

- **■** sample the phase space (2-dim)  $-1 < \cos \theta < 1$ ,  $0 < \phi < 2\pi$
- choosing  $\cos\theta$ ,  $\phi$  variables using uniformly distributed random number generator defines a candidate event
- lacktriangle d $\sigma$  is the event weight (probability of the event)
- $<\mathrm{d}\sigma>\sim\int\mathrm{d}\sigma$  converges to the cross section
- ullet at this point candidate events  $\theta\phi$  are distributed flat and carry no physics

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### Unweighting

If function to be integrated is a probability density (positive definite, f(x) > 0) one can convert it to arrive at a simulation of physical processes or Event Generator

- In addition to calculating the integral we often also want to select values of x (momenta,...) at random according to f(x). This is easy provided that we know the maximum value of the function in the region we are integrating over.
- lacksquare Then we randomly generate values of x in the integration region and keep them with probability

$$\mathcal{P} = \frac{f(x)}{f_{\text{max}}} \le R$$

- which is easy to implement by generating a random number between 0 and 1 and keeping the value of x if the random number R is less than the probability.
- This is called unweighting.

Selection of x according to f(x), in a random probabilistic way, event as they occur in Nature

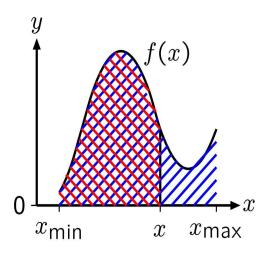
$$\int_{x_{min}}^{x} f(z)dz = R \int_{x_{min}}^{x_{max}} f(z)dz = RI$$

Selection of x according to f(x), in a random probabilistic way, event as they occur in Nature

$$\int_{x_{min}}^{x} f(z)dz = R \int_{x_{min}}^{x_{max}} f(z)dz = RI$$

Analytical (assumes primitive and its inverse known)

$$x = F^{-1}(F(x_{min}) + R I)$$



Selection of x according to f(x), in a random probabilistic way, event as they occur in Nature

$$\int_{x_{min}}^{x} f(z)dz = R \int_{x_{min}}^{x_{max}} f(z)dz = RI$$

ullet Hit and miss: assumes  $f_{max}$  known

$$I = \int_{x_{min}}^{x_{max}} f(x)dx = f_{max}(x_{max} - x_{min}) \frac{N_{acc}}{N_{tries}}$$
$$\frac{N_{acc}}{N_{tries}} = \text{efficiency}$$

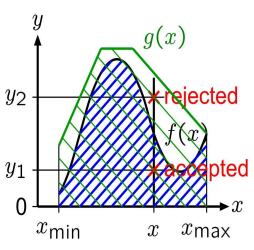
 $f_{\text{max}}$   $y_2$   $y_1$   $x_{\text{min}}$  x  $x_{\text{max}}$ 

MC → Event Generator, involves acceptance/rejection

Selection of x according to f(x), in a random probabilistic way, event as they occur in Nature

$$\int_{x_{min}}^{x} f(z)dz = R \int_{x_{min}}^{x_{max}} f(z)dz = RI$$

- $\bullet$  Importance Sampling: take f(x) < g(x) where G(x) and  $G^{-1}$  simple
- if y > f(x) go back to 1
- 1. select x according to g(x) 2. select y = Rg(x) (new R)



### Summary MC

Advantages of Monte Carlo Fast convergence in many dimensions

Arbitrarily complex integration regions

Few points needed to get first estimate

Each additional point improves the accuracy

Easy error estimate

More than one quantity can be evaluated at once.

- Disadvantages of Monte Carlo Slow convergence in few dimensions, but that is hardly the case in particle physics
- MC is well suited for particle physics where phase space integration involves a lot of variables with a complicated often not smooth function representing the cross section

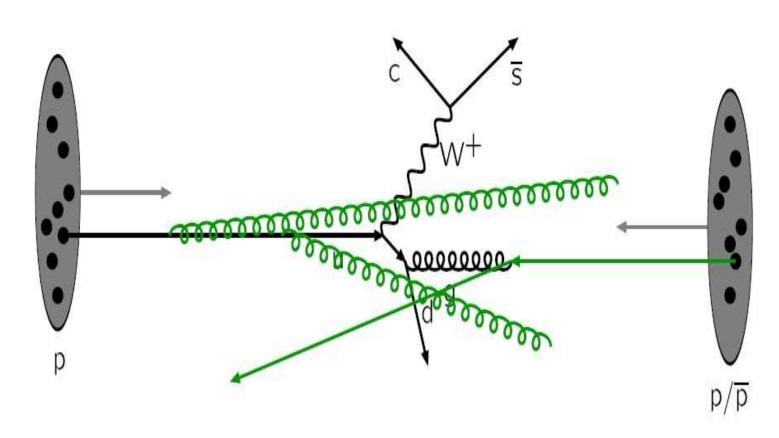
#### **Event Generator**

- With an integrand that is positive definite, which is the case for MC at LO, one deals with a probability. This lends itself to an event generator
- Allows a fully exclusive treatment exactly like real data
- At the most basic level a Monte Carlo event generator is a program which simulates particle physics events with the same probability as they occur in nature.
- In essence it performs a large number of integrals and then unweights to give the momenta of the particles which interact with the detector

# Event Generators, Radiation and Showers as (MC) probability



# Remember the Movie: The structure of an event, ISR and FSR

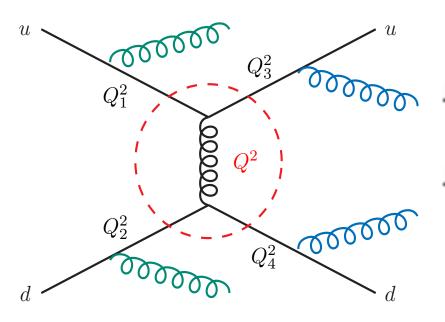


ISR: Initial State Radiation

# Parton Shower Approach

 $\mathcal{P}_{ ext{ISR}/ ext{FSR}}$  Accelerated charged particles radiate

$$2 \rightarrow n = (2 \rightarrow 2)_{\mathrm{On\ Shell}}$$
 + ISR + FSR

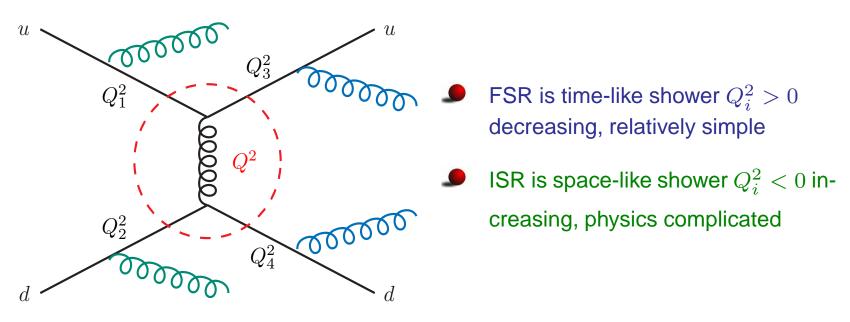


- FSR is time-like shower  $Q_i^2 > 0$  decreasing, relatively simple
  - ISR is space-like shower  $Q_i^2 < 0$  increasing, physics complicated

# Parton Shower Approach

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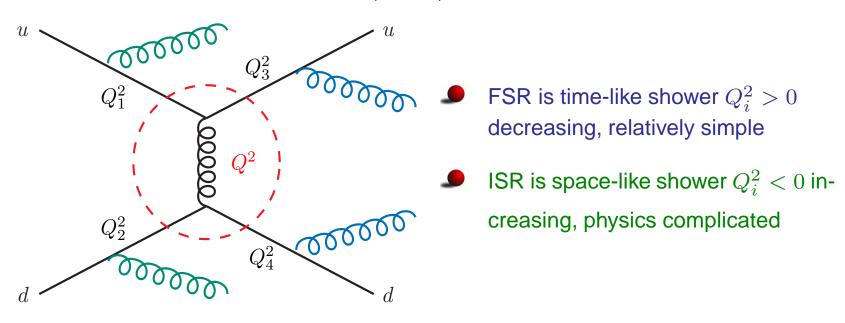


- Shower is viewed as a probabilistic process which occurs with unit probability
- The (total) cross section is not affected but indirectly it is since the event shape is changed
- Obviously it is an approximation to the full process

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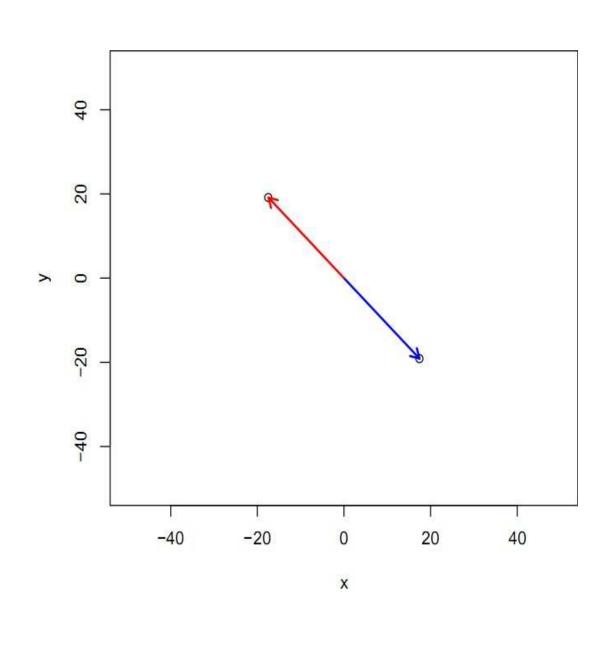
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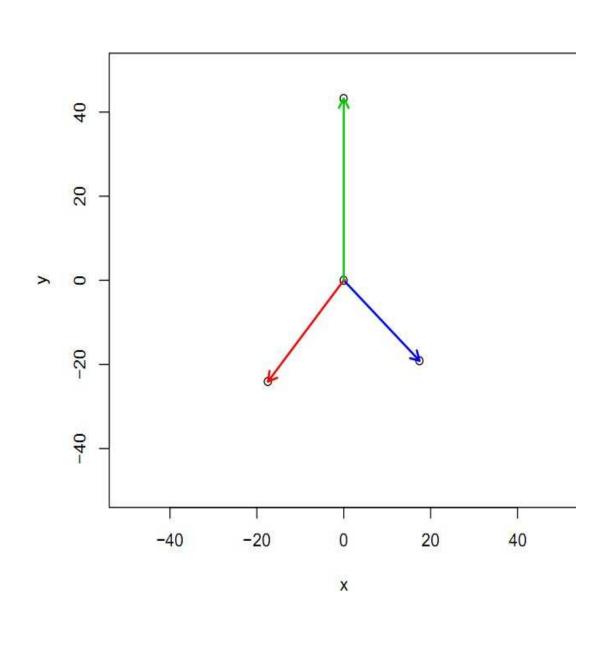


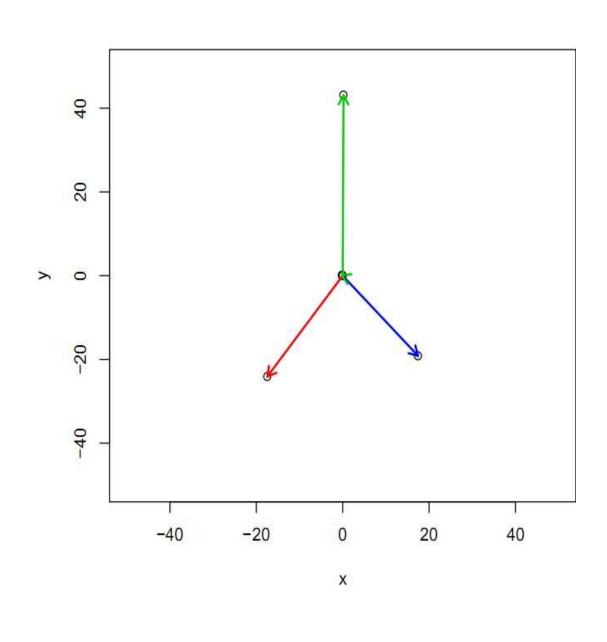
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- The (total) cross section is not affected but indirectly it is since the event shape is changed
- Obviously it is an approximation to the full process

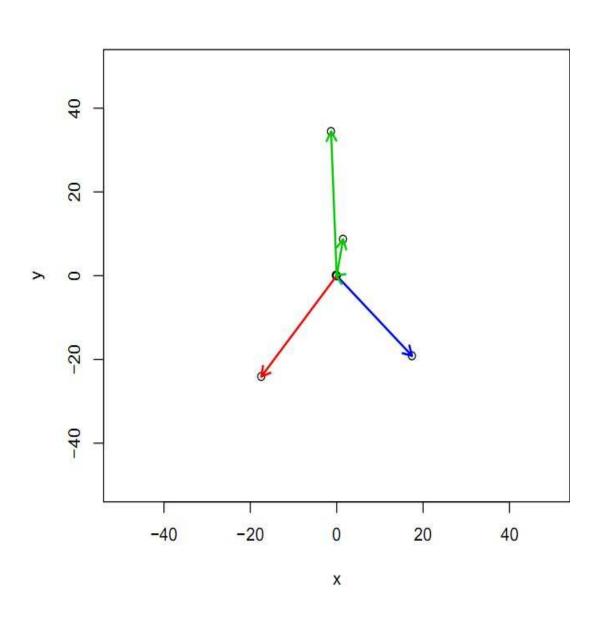


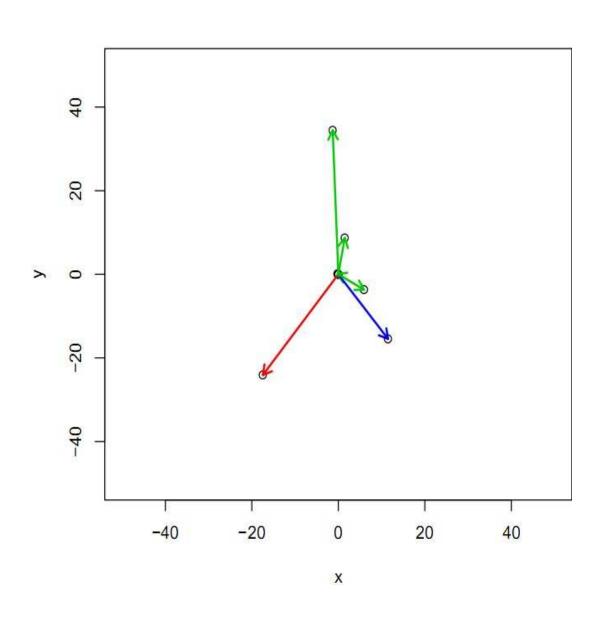
# Parton Shower movie

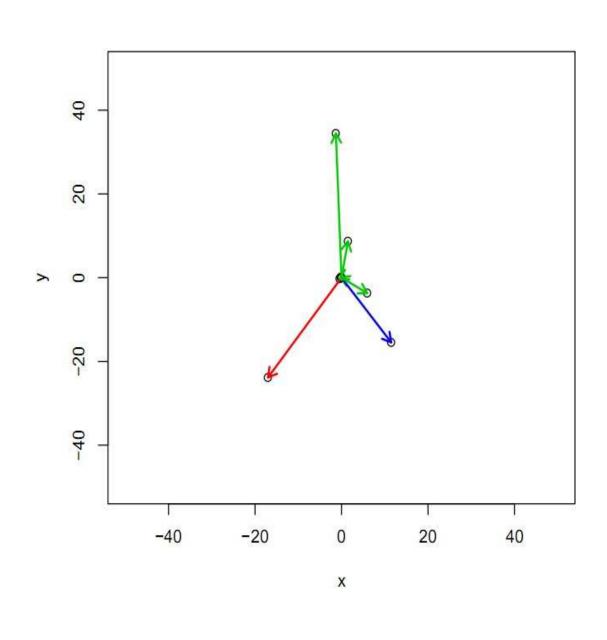


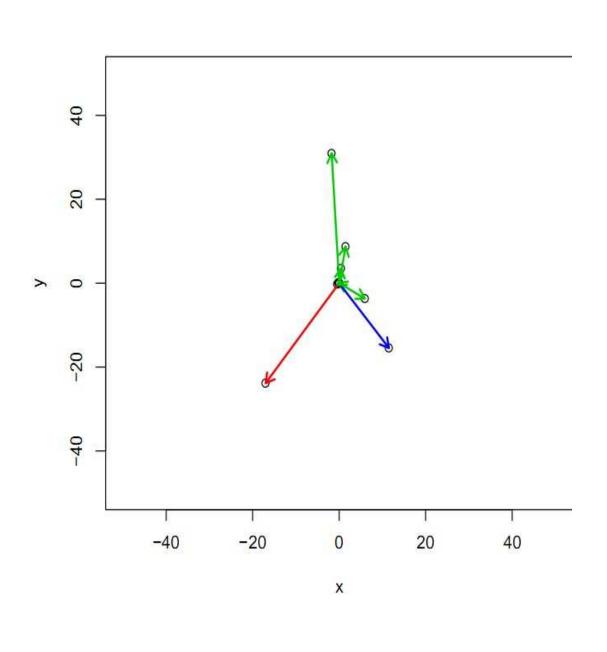


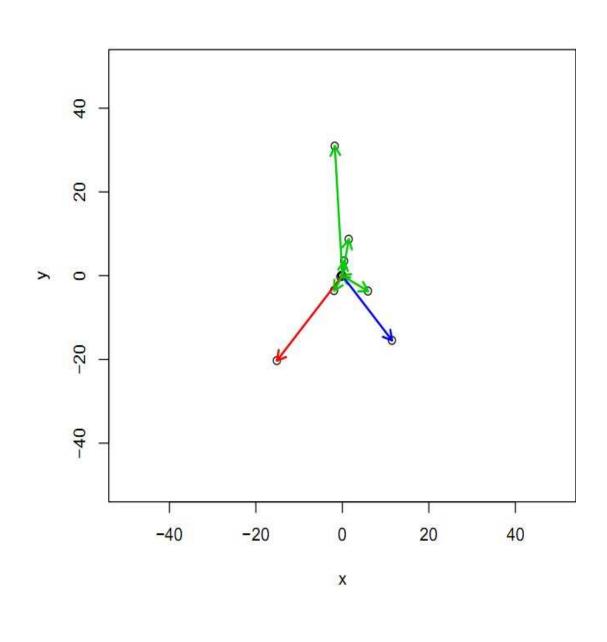


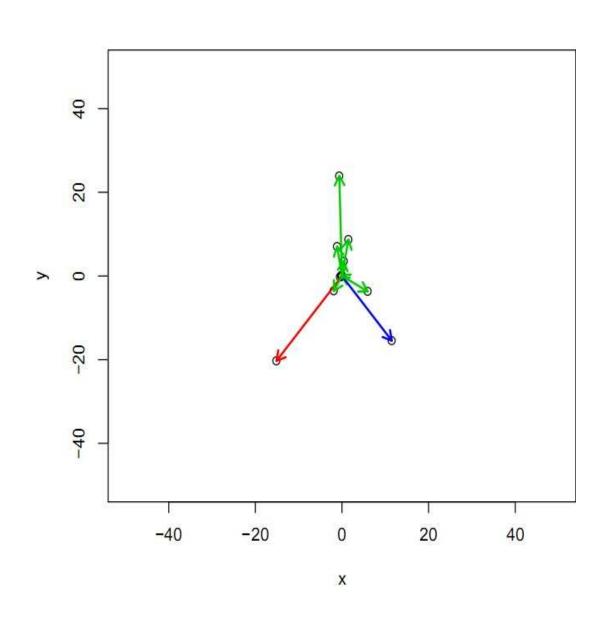


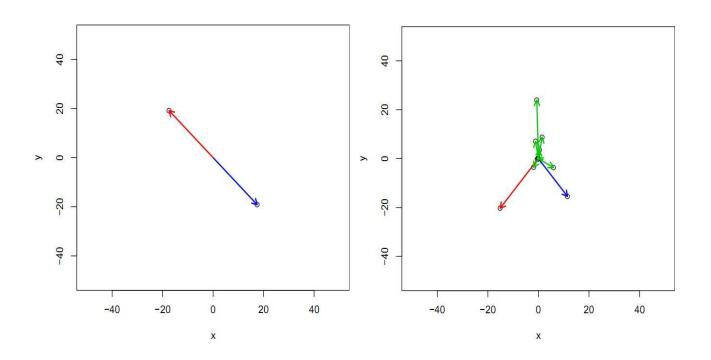




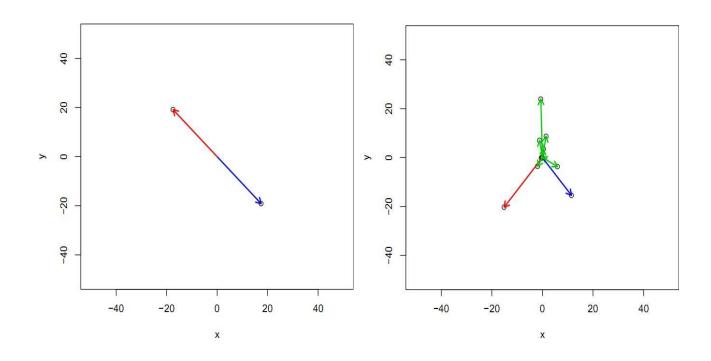




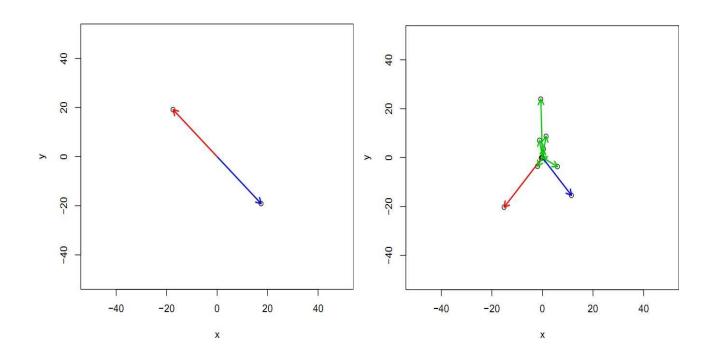




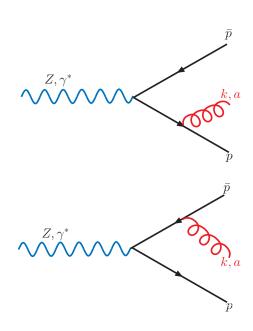
■ The topology generated by the PS can be quite complicated



- The topology generated by the PS can be quite complicated
- These are events shape that can not be described by fixed order pert. calculations



- The topology generated by the PS can be quite complicated
- These are events shape that can not be described by fixed order pert. calculations
- Total cross section still given by hard scattering (usually LO), experiments usually normalise to data

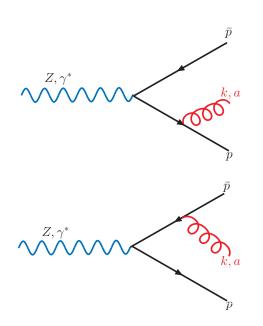


$$\mathcal{A}_{\mu} = \bar{u}(p) \not\in (-ig_s t_a) \frac{-i}{\not p' + \not k} \Gamma_{\mu} v(\bar{p}) \quad m_q = 0$$

$$+ \bar{u}(p) \Gamma_{\mu} \frac{i}{\not p' + \not k} (-ig_s t_a) \not\in v(\bar{p})$$

$$= -g_s \left( \frac{\bar{u}(p) \not\in (\not p' + \not k) \Gamma_{\mu} v(\bar{p})}{2p.k} - \frac{\bar{u}(p) \Gamma_{\mu} (\not p' + \not k) \not\in v(\bar{p})}{2\bar{p}.k} \right) t_a$$

$$\stackrel{Z,\gamma^*}{\sim} \qquad 2p.k = 4E_g E_p \sin^2 \left( \frac{\theta_{pk}}{2} \right) \rightarrow 0 \text{ for } E_g \rightarrow 0, \quad \theta_{pk} \rightarrow 0$$



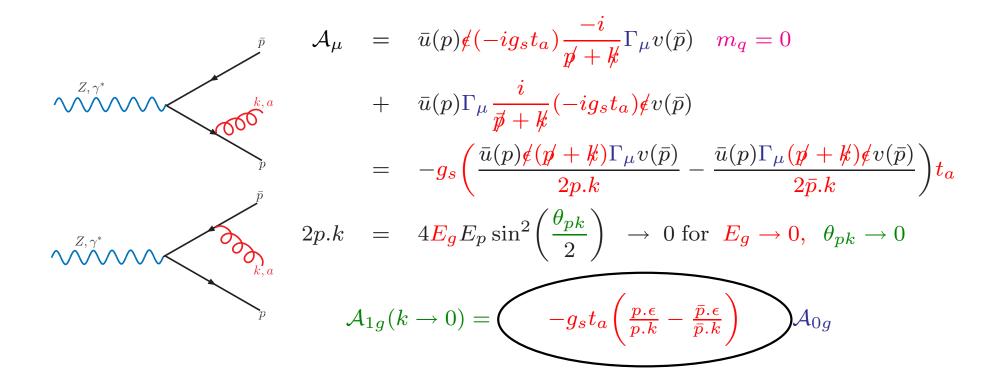
$$\mathcal{A}_{\mu} = \bar{u}(p) \not \in (-ig_s t_a) \frac{-i}{\not p + \not k} \Gamma_{\mu} v(\bar{p}) \quad m_q = 0$$

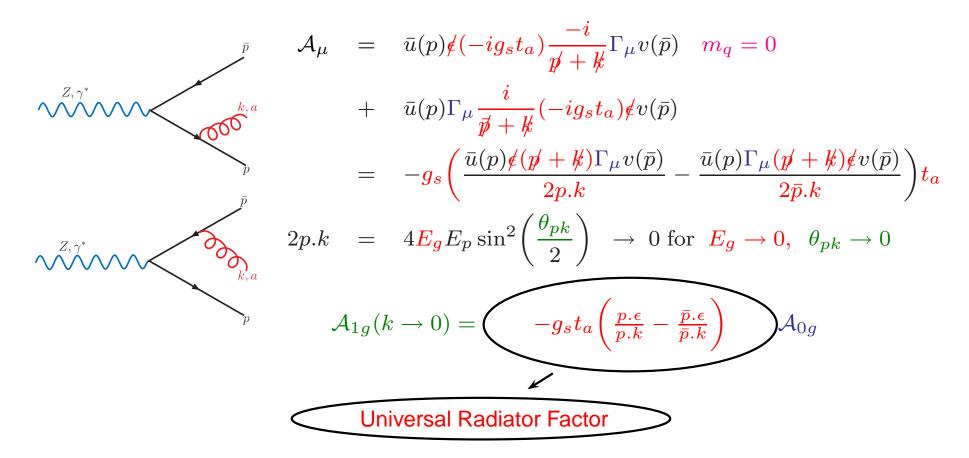
$$+ \bar{u}(p) \Gamma_{\mu} \frac{i}{\not p + \not k} (-ig_s t_a) \not \in v(\bar{p})$$

$$= -g_s \left( \frac{\bar{u}(p) \not \in (\not p + \not k) \Gamma_{\mu} v(\bar{p})}{2p.k} - \frac{\bar{u}(p) \Gamma_{\mu} (\not p + \not k) \not \in v(\bar{p})}{2\bar{p}.k} \right) t_a$$

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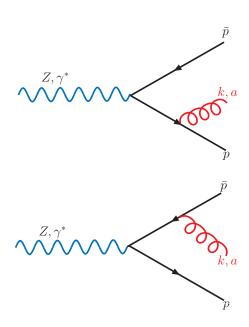
$$\mathcal{A}_{\text{soft}}(k \rightarrow 0) = -g_s t_a \left( \frac{p.\epsilon}{p.k} - \frac{\bar{p}.\epsilon}{\bar{p}.k} \right) \mathcal{A}_0 \quad \text{diverges } k \rightarrow 0$$

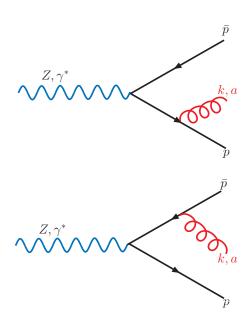




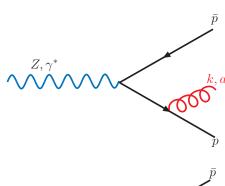
We have **factorisation** of the soft emission (long distance) from the short distance *i.e.* the **hard** process

## ${\sf Squaring\ soft/collinear}$



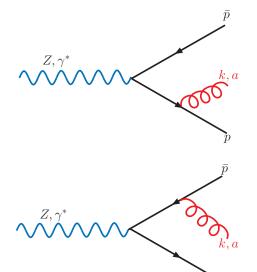


$$\mathcal{A}_{1g}(k \to 0) = -g_s t_a \left(\frac{p.\epsilon}{p.k} - \frac{\bar{p}.\epsilon}{\bar{p}.k}\right) \mathcal{A}_{0g}$$



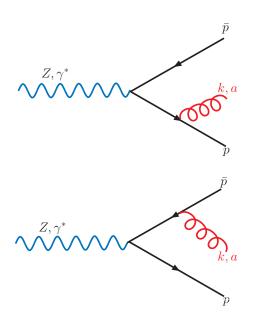
$$Z, \gamma^*$$
 $k, a$ 

$$\mathcal{A}_{1g}(k \to 0) = -g_s t_a \left(\frac{p.\epsilon}{p.k} - \frac{\bar{p}.\epsilon}{\bar{p}.k}\right) \mathcal{A}_{0g}$$
$$|\mathcal{M}_{1g}|^2 = \sum_{a,pol.(\epsilon)} |\mathcal{A}_{1g}(k \to 0)|^2 = C_F g_s^2 \frac{2p.\bar{p}}{p.k \; \bar{p}.k} |\mathcal{M}_{0g}|^2$$



$$\mathcal{A}_{1g}(k \to 0) = -g_s t_a \left(\frac{p.\epsilon}{p.k} - \frac{\bar{p}.\epsilon}{\bar{p}.k}\right) \mathcal{A}_{0g}$$

Phase Space 
$$|\mathcal{M}_{1g}|^2 d\Phi_{q\bar{q}g} = \left(|\mathcal{M}_{0g}|^2 d\Phi_{q\bar{q}}\right) d\mathcal{S}; d\mathcal{S} \simeq \frac{d^3\vec{k}}{2\omega_k (2\pi)^3} C_F g_s^2 \frac{2p.\bar{p}}{p.k\;\bar{p}.k}$$

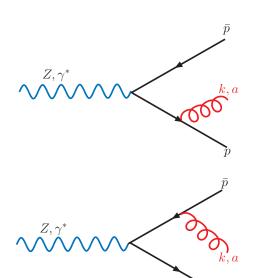


$$\mathcal{A}_{1g}(k \to 0) = -g_s t_a \left(\frac{p.\epsilon}{p.k} - \frac{\bar{p}.\epsilon}{\bar{p}.k}\right) \mathcal{A}_{0g}$$

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$$\theta = \theta_{\angle pk} , \ \phi = \text{azimuth}$$

$$d\mathcal{S} \simeq \frac{2\alpha_s C_F}{\pi} \frac{d\omega}{\omega} \frac{d\theta}{\sin\theta} \frac{d\phi}{2\pi}$$



$$\mathcal{A}_{1g}(k \to 0) = -g_s t_a \left(\frac{p.\epsilon}{p.k} - \frac{\bar{p}.\epsilon}{\bar{p}.k}\right) \mathcal{A}_{0g}$$

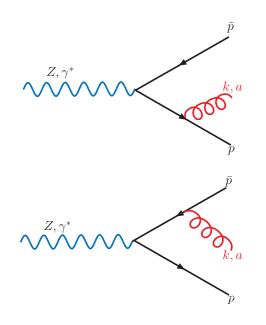
Phase Space 
$$|\mathcal{M}_{1g}|^2 d\Phi_{q\bar{q}g} = \left( |\mathcal{M}_{0g}|^2 d\Phi_{q\bar{q}} \right) d\mathcal{S}; d\mathcal{S} \simeq \frac{d^3\vec{k}}{2\omega_k (2\pi)^3} C_F g_s^2 \frac{2p.\bar{p}}{p.k \ \bar{p}.k}$$

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$$d\mathcal{S} \simeq \frac{2\alpha_s C_F}{\pi} \frac{d\omega}{\omega} \frac{d\theta}{\sin\theta} \frac{d\phi}{2\pi}$$

- $d\mathcal{S}$  diverges for  $\omega \to 0$ , Infrared divergence ( needs virtual loop corrections, we'll say more if time permits)
- $m extbf{ extit{9}} \ ext{d} \mathcal{S} ext{ diverges for } heta o 0 ext{ and } heta o \pi ext{ , collinear divergence}$

$$\mathcal{A}_{1g}(k \to 0) = -g_s t_a \left(\frac{p.\epsilon}{p.k} - \frac{\bar{p}.\epsilon}{\bar{p}.k}\right) \mathcal{A}_{0g}$$



Phase Space 
$$|\mathcal{M}_{1g}|^2 \mathrm{d}\Phi_{q\bar{q}g} = \left(|\mathcal{M}_{0g}|^2 \mathrm{d}\Phi_{q\bar{q}}\right) \mathrm{d}\mathcal{S}; \quad \mathrm{d}\mathcal{S} \simeq \frac{\mathrm{d}^3\vec{k}}{2\omega_k(2\pi)^3} C_F g_s^2 \frac{2p.\bar{p}}{p.k\,\bar{p}.k}$$

$$\theta = \theta_{\angle pk} \;, \; \phi = \mathrm{azimuth}$$

$$\mathrm{d}\mathcal{S} \simeq \frac{2\alpha_s C_F}{\pi} \frac{\mathrm{d}\omega}{\omega} \frac{\mathrm{d}\theta}{\sin\theta} \frac{\mathrm{d}\phi}{2\pi}$$

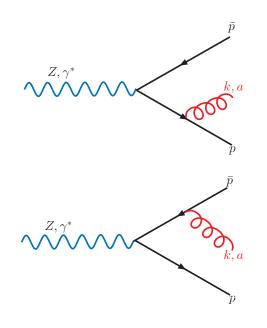
$$x_i = 2E_i/E_{\mathrm{tot}} \; p \to 1, \; k \to 3$$

$$\mathrm{d}\mathcal{S}_{\phi} = \frac{\alpha_s C_F}{2\pi} \mathrm{d}x_1 \mathrm{d}x_2 \frac{x_1^2 + x_2^2}{(1 - x_1)(1 - x_2)}$$

$$= \frac{\alpha_s C_F}{2\pi} \left(\frac{2}{\sin^2\theta} \frac{1 + (1 - x_3)^2}{x_2} - x_3\right) \mathrm{d}\cos\theta \mathrm{d}x_3$$

- $\bullet$  diverges for  $\omega \to 0$ , Infrared divergence (needs virtual loop corrections, we'll say more if time permits)
- lacksquare  $d\mathcal{S}$  diverges for heta o 0 and  $heta o\pi$  , collinear divergence

$$\mathcal{A}_{1g}(k \to 0) = -g_s t_a \left(\frac{p.\epsilon}{p.k} - \frac{\bar{p}.\epsilon}{\bar{p}.k}\right) \mathcal{A}_{0g}$$



Phase Space 
$$|\mathcal{M}_{1g}|^2 \mathrm{d}\Phi_{q\bar{q}g} = \left(|\mathcal{M}_{0g}|^2 \mathrm{d}\Phi_{q\bar{q}}\right) \mathrm{d}\mathcal{S}; \quad \mathrm{d}\mathcal{S} \simeq \frac{\mathrm{d}^3\vec{k}}{2\omega_k(2\pi)^3} C_F g_s^2 \frac{2p.\bar{p}}{p.k\,\bar{p}.k}$$

$$\theta = \theta_{\angle pk} \;, \; \phi = \mathrm{azimuth}$$

$$\mathrm{d}\mathcal{S} \simeq \frac{2\alpha_s C_F}{\pi} \frac{\mathrm{d}\omega}{\omega} \frac{\mathrm{d}\theta}{\sin\theta} \frac{\mathrm{d}\phi}{2\pi}$$

$$x_i = 2E_i/E_{\mathrm{tot}} \; p \to 1, \; k \to 3$$

$$\mathrm{d}\mathcal{S}_{\phi} = \frac{\alpha_s C_F}{2\pi} \mathrm{d}x_1 \mathrm{d}x_2 \frac{x_1^2 + x_2^2}{(1 - x_1)(1 - x_2)}$$

$$= \frac{\alpha_s C_F}{2\pi} \left(\frac{2}{\sin^2\theta} \frac{1 + (1 - x_3)^2}{x_2} - x_3\right) \mathrm{d}\cos\theta \mathrm{d}x_3$$

- $\mathcal{S}$  diverges for  $\omega \to 0$ , Infrared divergence (needs virtual loop corrections, we'll say more if time permits)
- $m extbf{ extit{9}} \ ext{d} \mathcal{S}$  diverges for heta o 0 and  $heta o \pi$  , collinear divergence
- ullet collinear divergence for  $x_1 o 1$  or  $x_2 o 1$  and Infrared divergence for  $x_3 o 0$

$$\frac{\mathrm{d}\mathcal{S}_{\phi}}{\sin^{2}\theta} \simeq \frac{\alpha_{s}C_{F}}{2\pi} \left( \frac{2}{\sin^{2}\theta} \frac{1 + (1 - x_{3})^{2}}{x_{3}} \right) \mathrm{d}\cos\theta \mathrm{d}x_{3}$$

$$\frac{2\mathrm{d}\cos\theta}{\sin^{2}\theta} = \frac{\mathrm{d}\cos\theta}{1 - \cos\theta} + \frac{\mathrm{d}\cos\theta}{1 + \cos\theta} = \frac{\mathrm{d}\cos\theta}{1 - \cos\theta} + \frac{\mathrm{d}\cos\bar{\theta}}{1 - \cos\bar{\theta}} \sim \frac{\mathrm{d}\theta^{2}}{\theta^{2}} + \frac{\mathrm{d}\bar{\theta}^{2}}{\bar{\theta}^{2}} \quad \text{for } \theta, \bar{\theta} \sim 0 \ll 1$$

$$\frac{\mathrm{d}\mathcal{S}_{\phi}}{\sin^{2}\theta} \simeq \frac{\alpha_{s}C_{F}}{2\pi} \left( \frac{2}{\sin^{2}\theta} \frac{1 + (1 - x_{3})^{2}}{x_{3}} \right) \mathrm{d}\cos\theta \mathrm{d}x_{3}$$

$$\frac{2\mathrm{d}\cos\theta}{\sin^{2}\theta} = \frac{\mathrm{d}\cos\theta}{1 - \cos\theta} + \frac{\mathrm{d}\cos\theta}{1 + \cos\theta} = \frac{\mathrm{d}\cos\theta}{1 - \cos\theta} + \frac{\mathrm{d}\cos\bar{\theta}}{1 - \cos\bar{\theta}} \sim \frac{\mathrm{d}\theta^{2}}{\theta^{2}} + \frac{\mathrm{d}\bar{\theta}^{2}}{\bar{\theta}^{2}} \quad \text{for } \theta, \bar{\theta} \sim 0 \ll 1$$

q and  $\bar{q}$  as independent emitters, notion of splitting as a probability

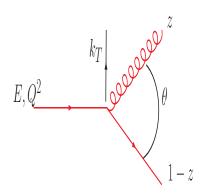
$$d\sigma_{1g} \sim d\sigma_{0g} \sum_{q \to qg}^{\bar{q} \to \bar{q}g} C_F \frac{\alpha_s}{2\pi} \frac{d\theta^2}{\theta^2} \frac{1 + (1-z)^2}{z} dz$$
  $(z \equiv x_3)$ 

$$\frac{\mathrm{d}\mathcal{S}_{\phi}}{\sin^{2}\theta} \simeq \frac{\alpha_{s}C_{F}}{2\pi} \left( \frac{2}{\sin^{2}\theta} \frac{1 + (1 - x_{3})^{2}}{x_{3}} \right) \mathrm{d}\cos\theta \mathrm{d}x_{3}$$

$$\frac{2\mathrm{d}\cos\theta}{\sin^{2}\theta} = \frac{\mathrm{d}\cos\theta}{1 - \cos\theta} + \frac{\mathrm{d}\cos\theta}{1 + \cos\theta} = \frac{\mathrm{d}\cos\theta}{1 - \cos\theta} + \frac{\mathrm{d}\cos\bar{\theta}}{1 - \cos\bar{\theta}} \sim \frac{\mathrm{d}\theta^{2}}{\theta^{2}} + \frac{\mathrm{d}\bar{\theta}^{2}}{\bar{\theta}^{2}} \quad \text{for } \theta, \bar{\theta} \sim 0 \ll 1$$

q and  $\bar{q}$  as independent emitters, notion of splitting as a probability

$$d\sigma_{1g} \sim d\sigma_{0g} \sum_{q \to qg}^{\bar{q} \to \bar{q}g} C_F \frac{\alpha_s}{2\pi} \frac{d\theta^2}{\theta^2} \frac{1 + (1 - z)^2}{z} dz \qquad (z \equiv x_3)$$

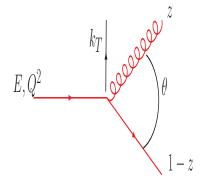


$$\frac{\mathrm{d}\mathcal{S}_{\phi}}{\sin^{2}\theta} \simeq \frac{\alpha_{s}C_{F}}{2\pi} \left( \frac{2}{\sin^{2}\theta} \frac{1 + (1 - x_{3})^{2}}{x_{3}} \right) \mathrm{d}\cos\theta \mathrm{d}x_{3}$$

$$\frac{2\mathrm{d}\cos\theta}{\sin^{2}\theta} = \frac{\mathrm{d}\cos\theta}{1 - \cos\theta} + \frac{\mathrm{d}\cos\theta}{1 + \cos\theta} = \frac{\mathrm{d}\cos\theta}{1 - \cos\theta} + \frac{\mathrm{d}\cos\bar{\theta}}{1 - \cos\bar{\theta}} \sim \frac{\mathrm{d}\theta^{2}}{\theta^{2}} + \frac{\mathrm{d}\bar{\theta}^{2}}{\bar{\theta}^{2}} \quad \text{for } \theta, \bar{\theta} \sim 0 \ll 1$$

q and  $\bar{q}$  as independent emitters, notion of splitting as a probability

$$d\sigma_{1g} \sim d\sigma_{0g} \sum_{q \to qg}^{\bar{q} \to \bar{q}g} C_F \frac{\alpha_s}{2\pi} \frac{d\theta^2}{\theta^2} \frac{1 + (1 - z)^2}{z} dz \qquad (z \equiv x_3)$$

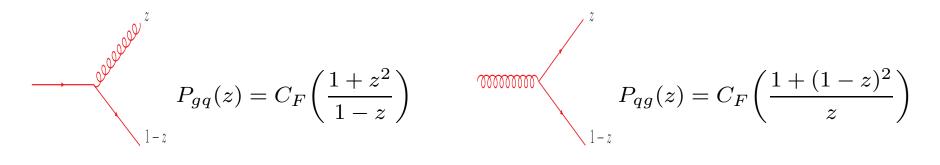


different choices of the evolution variables, equivalent in the limit (diff. in practice/different codes) 
$$Q^2 = E^2 z (1-z)\theta^2 \qquad k_T^2 = E^2 z^2 (1-z)^2 \theta^2$$
 
$$\frac{\mathrm{d}\theta^2}{\theta^2} = \frac{\mathrm{d}Q^2}{Q^2} = \frac{\mathrm{d}k_T^2}{k_T^2}$$

#### **DGLAP**

This generalises to different parton branching (gluon, quarks)

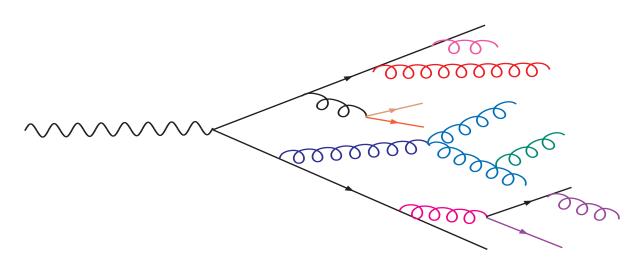
$$d\sigma_{bc} \sim d\sigma_a \frac{\alpha_s}{2\pi} \frac{d\theta^2}{\theta^2} P_{a \to bc}(z) dz$$



$$P_{qg}(z)=T_Rigg(z^2+(1-z)^2igg) \qquad T_R=rac{n_f}{2}$$
 (divergences at  $z=0,1$  dealt with soft/virtual corr.)

$$P_{gg}(z) = C_A \left( rac{z}{1-z} + rac{1-z}{z} + z(1-z) 
ight) \qquad C_A = 3 \qquad (C_F = 4/3)$$

Gluons radiate the most

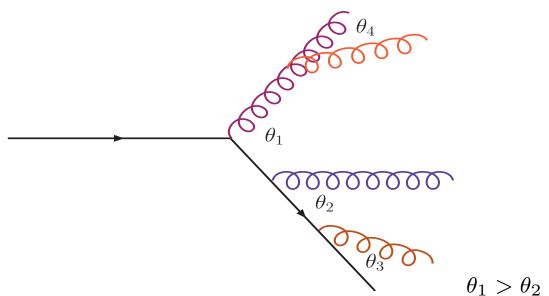


Need soft/collinear cut-offs to stay away from non perturbative physics. Details are model/code dependent

$$Q > m_0 = min(m_{ij}) \sim 1 \text{GeV}$$
  
 $z_{min}(E,Q) < z < z_{max}(E,Q)$ 

$$k_T > k_{T,min} \sim 0.5 {\rm GeV}$$

### Radiation is angle ordered



$$\theta_1 > \theta_2 > \theta_3$$
 and  $\theta_1 > \theta_4$ 

On average, emissions have decreasing angles with respect to emitters the jet is squeezed

#### The Probability of real emission exponentiates

- **Solution** Of total probability  $\mathcal{P}_{ ext{something}} + \mathcal{P}_{ ext{nothing}} = 1$ !
- Product of probabilities as time evolves  $T \sim 1/Q$  evolves

$$\mathcal{P}_{\text{nothing}}(0 < t < T) = \mathcal{P}_{\text{nothing}}(0 < t \le T_1)\mathcal{P}_{\text{nothing}}(T_1 < t \le T)$$

**subdivide further**  $T_i = (i/n)T, 0 \le i \le n$ 

$$\mathcal{P}_{\text{nothing}}(0 < t < T) = \lim_{n \to \infty} \prod_{i=0}^{n-1} \mathcal{P}_{\text{nothing}}(T_i < t < T_{i+1})$$

$$= \lim_{n \to \infty} \prod_{i=0}^{n-1} (1 - \mathcal{P}_{\text{something}}(T_i < t < T_{i+1}))$$

$$= \exp\left(-\lim_{n \to \infty} \sum_{i=0}^{n-1} \mathcal{P}_{\text{something}}(T_i < t < T_{i+1})\right)$$

$$= \exp\left(-\int_0^T \frac{d\mathcal{P}_{\text{something}}(t)}{dt} dt\right)$$

#### The Probability of real emission exponentiates

- Conservation of total probability  $\mathcal{P}_{\text{something}} + \mathcal{P}_{\text{nothing}} = 1$ !
- Product of probabilities as time evolves  $T \sim 1/Q$  evolves  $\mathcal{P}_{\text{nothing}}(0 < t < T) = \mathcal{P}_{\text{nothing}}(0 < t \le T_1)\mathcal{P}_{\text{nothing}}(T_1 < t \le T)$

**•** subdivide further 
$$T_i = (i/n)T, 0 \le i \le n$$

$$\mathcal{P}_{\text{nothing}}(0 < t < T) = \lim_{n \to \infty} \prod_{i=0}^{n-1} \mathcal{P}_{\text{nothing}}(T_i < t < T_{i+1})$$

$$= \lim_{n \to \infty} \prod_{i=0}^{n-1} (1 - \mathcal{P}_{\text{something}}(T_i < t < T_{i+1}))$$

$$= \exp\left(-\lim_{n \to \infty} \sum_{i=0}^{n-1} \mathcal{P}_{\text{something}}(T_i < t < T_{i+1})\right)$$

$$= \exp\left(-\int_0^T \frac{d\mathcal{P}_{\text{something}}(t)}{dt} dt\right)$$

$$d\mathcal{P}_{\text{first}}(T) = d\mathcal{P}_{\text{something}}(T) \exp\left(-\int_{0}^{T} \frac{d\mathcal{P}_{\text{something}}(t)}{dt} dt\right)$$

### The Probability of real emission exponentiates

$$d\mathcal{P}_{\text{first}}(T) = d\mathcal{P}_{\text{something}}(T) \exp\left(-\int_{0}^{T} \frac{d\mathcal{P}_{\text{something}}(t)}{dt} dt\right)$$

#### The Probability of real emission exponentiates

$$d\mathcal{P}_{\text{first}}(T) = d\mathcal{P}_{\text{something}}(T) \exp\left(-\int_0^T \frac{d\mathcal{P}_{\text{something}}(t)}{dt}dt\right)$$

$$d\mathcal{P}_{a\to bc} = \frac{\alpha_s}{2\pi} \frac{dQ^2}{Q^2} P_{a\to bc}(z) dz \left( \exp\left(-\sum_{bc} \int_{Q^2}^{Q_{\text{max}}^2} \frac{dQ'^2}{Q'^2} \int_{Q_0^2/Q^2}^{1-Q_0^2/Q^2} \frac{\alpha_s}{2\pi} P_{a\to bc}(z') dz' \right) \right)$$

 $Q_0 = \text{low cut-off scale}$ 

#### The Probability of real emission exponentiates

$$d\mathcal{P}_{\text{first}}(T) = d\mathcal{P}_{\text{something}}(T) \exp\left(-\int_0^T \frac{d\mathcal{P}_{\text{something}}(t)}{dt}dt\right)$$

$$d\mathcal{P}_{a\to bc} = \frac{\alpha_s}{2\pi} \frac{dQ^2}{Q^2} P_{a\to bc}(z) dz \left( \exp\left(-\sum_{bc} \int_{Q^2}^{Q_{\max}^2} \frac{dQ'^2}{Q'^2} \int_{Q_0^2/Q^2}^{1-Q_0^2/Q^2} \frac{\alpha_s}{2\pi} P_{a\to bc}(z') dz' \right) \right)$$

 $Q_0 = low cut-off scale$ 

 $\Delta(Q^2,Q^2_{
m max})$ , Sudakov form factor (probability of emitting no radiation between these 2 scales)  ${\cal P}_{
m nothing}$ 

(a given parton only branches once)

#### Numerical MC Procedure of PS

- Start with a parton at high  $Q_{max}^2$  (typical of hard process)
- Work out the scale of the next branching,  $Q^2$  by generating a random number  $R\in[0,1]$  and solving  $R=\Delta(Q^2_{max},Q^2)$
- lacksquare if no solution  $Q^2>Q_0^2$  stop
- otherwise work out the type of the branching
- generate the momenta of the decay products using the splitting functions
- repeat the procedure for the newly produced partons

#### (some) differences between the MC for PS

key difference is the evolution/scale variable

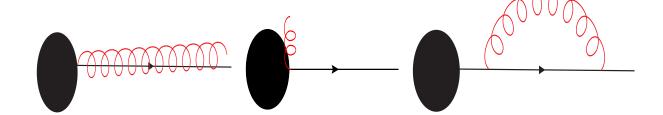
Angle  $\theta$  (ordering HERWIG

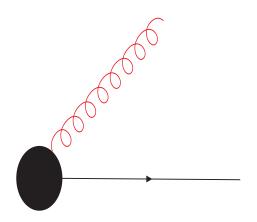
Virtuality  $Q^2$ 

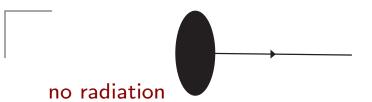
Transverse momentum  $k_T$ 

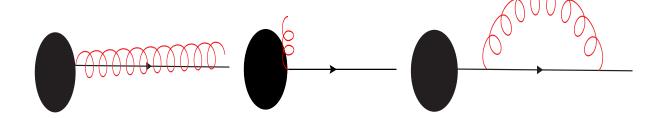
- soft emission (coherence), recall this factorises at the amplitude level...

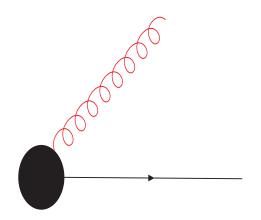




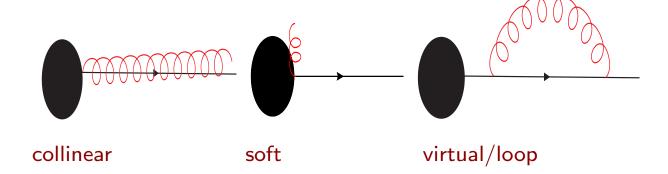


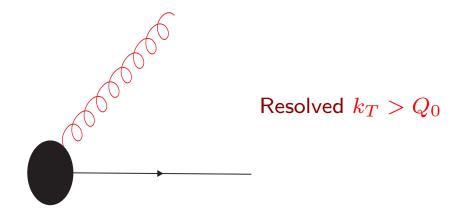


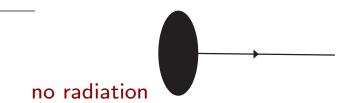


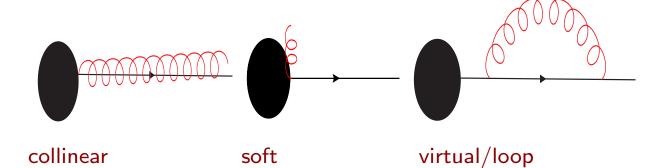




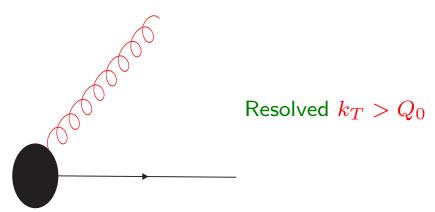








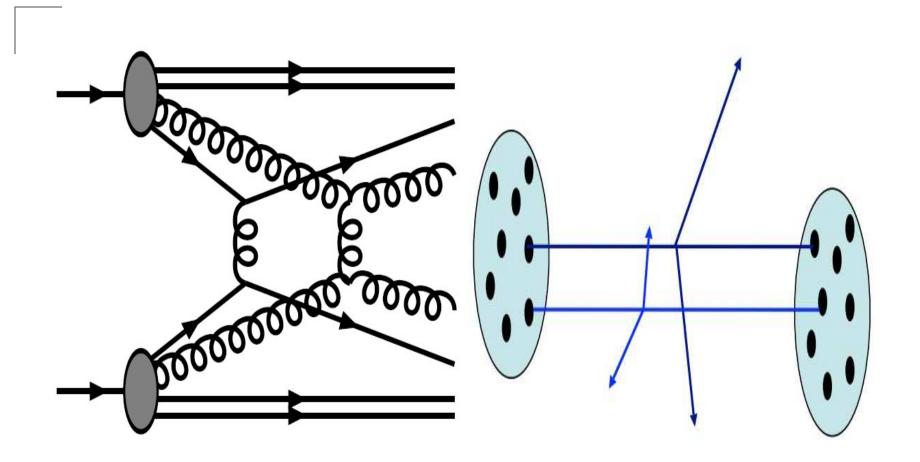
Unresolved from no radiation,  $k_T < Q_0$ . With addition of virtual, divergence tamed



## Event Generators, other (non perturbative stuff)



## Multiple Parton Interaction



#### Multiple Parton Interaction

- $m Partial P_{T,\min}$  and high energy inclusive parton-parton cross section is larger than proton-proton cross section
- More than one parton (per proton) scatter
- calls for a model of spatial distribution within the proton (perturbation theory gives n-scatter distribution
- UE is not understood. It is modelled through some data that do not give strong constraints and hence extrapolation can be doubted.(minimum bias)

